Geological Society of America Special Papers Online First

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Geological Society of America Special Papers, published online July 22, 2014; doi:10.1130/2014.2507(01)

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THE GEOLOGICAL SOCIETY OF AMERICA The Geological Society of America Special Paper 507 2014

Early Cretaceous to present latitude of the central proto-Tibetan Plateau: A paleomagnetic synthesis with implications for Cenozoic tectonics, paleogeography, and climate of Asia

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ABSTRACT

Published paleomagnetic data from well-dated sedimentary rocks and lavas from the Lhasa terrane have been reevaluated in a statistically consistent framework to assess the latitude history of southern Tibet from ca. 110 Ma to the present. The resulting apparent polar wander path shows that the margin of the Lhasa terrane has remained at lat $\sim 20^\circ \pm 4^\circ N$ from ca. 110 to at least 50 Ma and has drifted northward to its present latitude of 29°N since the early Eocene. This latitude history provides a paleomagnetically determined collision age between the Tibetan Himalaya and the southern margin of Asia that is ca. 49.5 ± 4.5 Ma, if not a few millions of years earlier after considering reasonable estimates for shortening within the suture zone. This collision occurred at lat $\sim 21^{\circ} \pm 4^{\circ}$ N, or perhaps $\sim 2^{\circ}$ lower if an average-size forearc is considered. These paleomagnetic data indicate that at most, only 1100 ± 560 km of post-50 Ma India-Asia convergence was partitioned into Asian lithosphere. The lower bound of these paleomagnetic estimates is consistent with the magnitude of upper crustal shortening and thickening within Asia calculated from structural geologic studies. Thus, a substantial amount of the shortening within, and therefore surface uplift of, the Tibetan Plateau predates the Tibetan Himalaya-Lhasa collision. These conclusions suggest that the Tibetan Plateau is similar to the Altiplano of the Andes,

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Lippert, P.C., van Hinsbergen, D.J.J., and Dupont-Nivet, G., 2014, Early Cretaceous to present latitude of the central proto-Tibetan Plateau: A paleomagnetic synthesis with implications for Cenozoic tectonics, paleogeography, and climate of Asia, *in* Nie, J., Horton, B.K., and Hoke, G.D., eds., Toward an Improved Understanding of Uplift Mechanisms and the Elevation History of the Tibetan Plateau: Geological Society of America Special Paper 507, doi:10.1130/2014.2507(01). For permission to copy, contact editing@geosociety.org. © 2014 The Geological Society of America. All rights reserved.

in that most of the plateau developed at subtropical latitudes above an oceanic subduction zone in the absence of a continent-continent collision. A direct implication of these findings is that 1700 ± 560 km or more post–50 Ma India-Asia convergence was partitioned into the lower plate of the orogenic system (i.e., units of Indian affinity). Recent paleomagnetic and plate tectonic analyses suggested significant extension of Greater India lithosphere after breakup from Gondwana but prior to collision with the southern margin of Asia. Cretaceous extension within Greater India was inferred to open an oceanic Greater India Basin, which would have maintained a deep tropical water mass along the southern edge of greater Asia throughout most of the Paleogene. We suggest ways in which future climate models can incorporate this paleogeography to more accurately explore how Paleogene atmospheric processes interact with or are modified by the juxtaposition of a tropical ocean basin and the high uniform topography of the Tibetan Plateau.

INTRODUCTION

The thick crust and elevated, low relief surfaces of orogenic plateaus such as the Altiplano in the Andes, the Turkish-Iranian high plateau, or the Tibetan Plateau are among the most conspicuous physiographic features on Earth. These first-order attributes of plateaus have motivated considerable research exploring possible feedbacks between plateau growth and plate force balances (e.g., for Tibet, Molnar and Lyon-Caen, 1988; Flesch et al., 2005; Iaffaldano et al., 2007; Molnar and Stock, 2009; Copley et al., 2010, 2011; for the Andes, Oncken et al., 2006; for the Turkish-Iranian plateau, Allen et al., 2004; Sengör et al., 2008), as well as potential feedbacks between plateau growth and atmospheric dynamics and related climate evolution (e.g., Manabe and Terpstra, 1974; Sobel et al., 2003; see review in Molnar et al., 2010). In general, plateaus result from distributed compressional deformation at leading edges of overriding continental plates above subduction zones, followed, in places, by continent-continent collision. The Tibetan Plateau is an ideal case study for understanding partitioning of strain related to India-Asia convergence into the overriding Asian plate since the Early Cretaceous, first during oceanic subduction and subsequently during continent-continent collision stages (Murphy et al., 1997; DeCelles et al., 2007; Kapp et al., 2007a, 2007b; Rohrmann et al., 2012). This strain is now manifested as the modern Tibetan Plateau, which creates a barrier to west-to-east subtropical tropospheric air flow and seasonally pulls the Intertropical Convergence Zone (ITCZ) far north of the equator. These perturbations to atmospheric circulation affect not just the regional climate of Asia, but also possibly the global distribution of heat and moisture (see reviews in Rodwell and Hoskins, 1996; Molnar et al., 2010; Huber and Goldner, 2012). The plateau and high Himalaya also prevent cool, dry subtropical air from Central Asia from mixing with warm, humid tropical air from the Indian Ocean, and therefore may simultaneously accentuate the Indian monsoon and the subtropical aridity that characterizes the climate of western China and surrounding regions (Boos and Kuang, 2010).

Similar processes probably have been active since the plateau first reached an as-yet unknown critical extent and height, but the response of atmospheric dynamics and feedbacks with surface processes to the presence of a Tibetan Plateau also may be sensitive to its latitudinal position (e.g., within the arid subtropics; Sobel et al., 2003). Thus, climate models that explore the role of the Tibetan Plateau in regional climate change should evaluate the distribution of high-elevation, low-relief topography not only in time, but also in (latitudinal) space. A key first step in addressing any of these relationships is to quantitatively constrain the paleogeography of the Tibetan Plateau throughout its geologic history in terms of its surface uplift, its structural position relative to Eurasia and India, and its latitude. In this paper we address the latter history using paleomagnetism. The convergence between India and Asia is essentially north-south, and positions of the Indian and Asian continents are reconstructed in detail based on marine magnetic anomalies (e.g., Molnar and Stock, 2009; Copley et al., 2010; Cande and Stegman, 2011; van Hinsbergen et al., 2011b). Paleomagnetic data, and paleolatitudinal motion derived from inclination values, are thus very useful tools to date collisional events from rocks of the deformed northern Indian and southern Eurasian margin, and can quantify how convergence is partitioned into strain in the upper and lower plates of the orogen (Besse et al., 1984; Chen et al., 1993; Dupont-Nivet et al., 2010a; Lippert et al., 2011; van Hinsbergen et al., 2012b).

Several recent paleomagnetic studies of volcanic and sedimentary units of Cretaceous to Paleogene age from the Lhasa terrane allow us to construct a detailed paleolatitude record for southern Tibet. The results from these studies are highly variable, however, and have been used at face value to support disparate paleogeographic reconstructions. For example, Achache et al. (1984), Tan et al. (2010), Dupont-Nivet et al. (2010a), Liebke et al. (2010), Sun et al. (2010), and Chen et al. (2010) each sampled the upper Linzizong Formation (Pana formation of Liu, 1993), which consists of felsic volcanic rocks and interlayered terrestrial sediments and is widely distributed across the southern Lhasa terrane, and has been well dated to 57-46 Ma using both ⁴⁰Ar/³⁹Ar (Lee et al., 2009) and U-Pb (He et al., 2007) isotopic systems. Calculated paleolatitudes range from 8°N (Chen et al., 2010) to 33°N (Tan et al., 2010). It is not possible for all of these individual paleolatitude estimates to be accurate within such a short time interval.

The purpose of this study is to utilize the wealth of new paleomagnetic data from the Lhasa terrane to produce a robust paleolatitude history of the southern margin of the Tibetan Plateau. We accomplish this by combining and interpreting timeequivalent paleomagnetic data in a coherent and statistically equivalent framework. Previous reviews of paleomagnetic data from the Lhasa terrane (e.g., Ali and Aitchison, 2006; Liebke et al., 2010; Najman et al., 2010) have treated published averaged paleomagnetic data with equal statistical weight even though these averaged data are disparate in terms of sample size and underlying selection criteria. Herein we demonstrate why this approach should be used with caution in the context of the India-Asia collision. Our approach is strictly a statistical one, with an emphasis on applying the same methodology to each data set to establish coherency within and between paleomagnetic data sets from southern Tibet. Our synthetic approach should provide data sets that are large enough and varied enough in sampling strategy to perform statistically robust and geologically meaningful paleomagnetic field tests and tests for averaging of paleosecular variation in the cases involving volcanic units. Our approach should also help identify the source or cause of the variability between the results from individual paleomagnetic studies, indicate sampling gaps in the paleomagnetic record, and provide a latitude history for southern Tibet that can be tested by future research. This information is crucial for designing sampling strategies for field campaigns that will maximize the utility, quality, and robustness of future paleomagnetic studies of the Lhasa terrane.

We summarize the published paleomagnetic data from Cretaceous to Paleogene stratigraphic units from the Lhasa terrane and then evaluate these data for consistent quality using the most up-to-date criteria achieved by the geomagnetic field community. This study is an update of paleomagnetic data presented in van Hinsbergen et al. (2012b, their supplementary information), and here is discussed specifically in the context of the Lhasa terrane and Tibetan Plateau. We discuss these results in terms of calculating the paleolatitude for the southern margin of Asia since ca. 110 Ma. We conclude with a brief discussion of implications for the Cenozoic tectonics of the India-Asia collision and Tibetan Plateau and regional climate dynamics; we also highlight potential directions for future research.

METHODS: PALEOMAGNETIC ARCHIVES OF LATITUDE

Global Synthetic Apparent Polar Wander Path and Kinematic Plate Reconstructions: Positioning India and Asia

The relative positions of plates that are bounded by active or former mid-ocean ridges are determined by plate circuits based on marine magnetic anomalies and fracture zones (Cox and Hart, 1986). The position of the plate circuit is constrained relative to the dipolar geomagnetic field assumed to parallel the Earth's spin axis in a so-called global apparent polar wander path (GAPWP), which utilizes paleomagnetic data from the stable continents. Here we use the most recent GAPWP of Torsvik et al. (2012), which was constructed by first rotating all available high-quality paleomagnetic poles from each stable continental region to a common reference location (e.g., South African plate) using the Euler poles from the plate circuit, then time-averaging these paleomagnetic data, and finally transforming this smoothed paleomagnetic pole path back to each continental reference frame using the same plate circuit.

The GAPWP provides paleolatitudes to within a few degrees uncertainty, because the synthetic nature and long-time averaging of synthetic GAPWPs reduce uncertainties and introduce some smoothing in the position of the mean apparent paleomagnetic pole for a continental block (for a review, see Torsvik et al., 2012). It is important to note that some Paleogene paleomagnetic data from Central and East Asia yield anomalously low paleolatitudes relative to the global database (e.g., Si and van der Voo, 2001; Cogné et al., 1999, 2013; Dupont-Nivet et al., 2010b; Lippert et al., 2011). These data were omitted from the GAPWP compilation of Torsvik et al. (2012). The GAPWP provides precise positions of stable Siberia-Eurasia and India.

Despite their precision, plate circuits provide only relative positions and motions of one plate relative to another, and the GAPWP provides precise locations only of the stable regions of continents relative to the spin axis. Relics of former plates such as the Tibetan microcontinental fragments or intensely deformed Indian and Asian plate boundaries cannot be directly reconstructed in a plate circuit and GAPWP, contrary to some interpretations (Ali and Aitchison, 2004, 2006; Aitchison et al., 2007). The only way to quantitatively place such regions, including Tibet, into past plate paleogeographic configurations is by using paleomagnetic data, constraining paleolatitude and vertical axis rotation (but not paleolongitude). Thus, placing highly deformed regions like southern Tibet within the context of the global plate circuit, i.e., quantifying the paleogeography of orogens in space and time, requires paleomagnetic data from the orogen. The paleolatitudes of deformed continental margins like the northern edge of Greater India and the southern edge of the Tibetan orogen, i.e., zero-order paleogeographic boundary conditions for calculating continental collision events, the distribution of strain within the orogen, and feedbacks between topography, atmospheric dynamics, and climate, must be calculated from robust paleomagnetic data from the margins (e.g., Dupont-Nivet et al., 2010a; Yi et al., 2011).

Positioning Tibet within India-Asia Convergence

Data Selection

A robust paleomagnetic pole is calculated from primary, welldated, well-determined, independent readings of the geomagnetic field over a time period that is long enough to average the effect of secular variation, but short enough so as to not integrate significant plate motion (e.g., van der Voo, 1990). This record can come from sedimentary rocks and lavas, although evaluating the fidelity of this record is very different for the two media. It is generally

assumed that continuous sections of sedimentary rocks naturally average secular variation, such that a stratigraphic series of site mean directions or a magnetostratigraphic column will provide a good characterization of the magnetic field. This assumption is only valid, however, if the preserved magnetic directions are primary and have not been biased by sedimentary inclination shallowing caused by particle settling and compaction (King, 1955). Sedimentary inclination shallowing is especially prevalent in detrital rocks (Lovlie and Torsvik, 1984; Tauxe and Kent, 2004; Kodama, 2009) and is well documented throughout Asia (e.g., Gilder et al., 2001; Yan et al., 2005; see also recent reviews in Dupont-Nivet et al., 2010b; Lippert et al., 2011). Therefore, paleomagnetic results from sediments should be used for paleolatitude reconstructions only after they have been evaluated for this inclination shallowing bias. The likelihood that a sedimentary paleomagnetic data set represents a time-averaged geomagnetic field direction can be evaluated with reasonable assumptions about geomagnetic field behavior (Deenen et al., 2011).

In principle, volcanic rocks are immune to the inclination shallowing bias that appears prevalent in many detrital rocks. Because the processes and time scales by which and over which magnetic directions are recorded in volcanic rocks are fundamentally different from those in sediments, using the paleomagnetism of lavas for tectonic investigations poses unique sampling requirements. Cooling of a single, thin (1-3 m) lava unit will record a geologically instantaneous measurement of the geomagnetic field (~3–30 days; Turcotte and Schubert, 2002). Furthermore, lavas can extrude rapidly as well as irregularly, such that several stratigraphically successive lavas often record the same magnetic field direction. Therefore, paleomagnetic data from volcanic units should first be filtered so that all site mean directions are of comparably high statistical quality, and then, if the sampling strategy permits, stratigraphically successive site mean directions that are statistically indistinguishable should be combined into direction groups (e.g., Mankinen et al., 1985). This procedure ensures that several independent readings of the geomagnetic field have been sampled to accurately characterize its time-averaged behavior and direction.

In this study, we apply the following filtering criteria to site mean directions determined from lavas by excluding sites that (1) are not used by the original authors if reason for exclusion is provided; (2) contain directions of mixed polarity, as lava sites should be spot readings that cannot record a reversal; (3) are characterized by fewer than five samples; (4) have k-values (Fisher's [1953] precision parameter) <50; or (5) or are beyond the angular threshold determined by the Vandamme (1994) cut-off criteria.

This filtering procedure is commonly used by the geomagnetic field community (Johnson et al., 2008), but we note that this treatment of volcanic-based paleomagnetic data is seldom employed in the tectonics community.

Data Processing

Paleomagnetic studies of the present and ancient magnetic field indicate that virtual geomagnetic poles (VGPs) fit a Fisherian distribution better than the parent directions (e.g., Johnson et al., 2008). Therefore, we transform paleomagnetic directions into VGPs and calculate the Fisher mean VGP for each sampling locality. We note that use of the Vandamme (1994) cut-off criteria or a fixed 45° angular cut-off for identifying directional outliers provides the same results for the data sets examined in the following. The assignment of direction groups follows from the original authors of each of the studies. We evaluate the likelihood that a mean lava-based VGP represents the time-averaged geomagnetic field following Deenen et al. (2011). If the calculated error (A95) of the mean VGP is within the error envelope that can be expected based on paleosecular variation models, expressed as A95_{min} and A95_{max} (minimum and maximum), then we conclude that the mean VGP is a time-averaged paleomagnetic pole suitable for paleogeographic studies.

This approach is similar to testing if the angular dispersion (S) of VGPs is similar to the dispersion observed in ancient lavas that yield good estimates of a geocentric axial dipole (GAD) field when averaged (Biggin et al., 2008; Johnson et al., 2008) and to the dispersion determined from paleosecular variation models like TK03.GAD (Tauxe and Kent, 2004). Although the use of S for evaluating time averaging of secular variation has been disputed (Linder and Gilder, 2012), we contend that direct comparison of the directional distributions of ancient lava-based paleomagnetic data to those of large sets of recent lava-based paleomagnetic data (in contrast to comparisons to numerical models or sedimentary data) is the most reasonable approach. Notably, if the angular dispersion of the VGPs is consistent with dispersion observed in young lava successions, then we conclude that the sample set characterizes the time-averaged geomagnetic field.

To determine the Cretaceous-Paleogene latitude of the southern margin of the Lhasa terrane, we calculate the expected paleolatitude and its uncertainty at a reference location on the Indus-Yarlung suture zone (IYSZ, 29°N, 88°E) from each accepted paleomagnetic pole using equations provided in Butler (1992). Until the ca. 55-50 Ma arrival of the Tibetan Himalaya in the subduction zone below Tibet, the southern margin of Asia was located along the modern IYSZ that generally separates rocks of Asian affinity from those of Tibetan Himalayan affinity (IYSZ, Fig. 1) (Hodges, 2000; Hébert et al., 2012). Tibetan Himalayan upper crustal rocks subsequently accreted to and deformed southern Asia, and the southern margin of Asia consequently migrated south of the IYSZ (van Hinsbergen et al., 2012a, 2012b). To keep the same reference point through time, we identify the location of the southern Asian plate boundary until 50 Ma, and the Indus-Yarlung suture thereafter. All calculations are made assuming a 100% axial dipolar geomagnetic field (cf. Si and van der Voo, 2001; Dupont-Nivet et al., 2010b). Paleomagnetic poles and paleolatitudes are listed in Table 1. All age assignments follow the original authors unless otherwise noted and are keyed to the same time scales as used in the global synthetic GAPWP and kinematic plate reconstructions (Torsvik et al., 2012); i.e., Gradstein et al. (1994) for the Mesozoic, and Cande and Kent (1995) for the Late Cretaceous and younger, unless the section is dated using isotopic age determinations.





Figure 1. Tectonostratigraphic map of the central Lhasa terrane, southern Tibet (modified by Kapp et al., 2007b). Numbers next to blue circles correspond to locations of paleomagnetic sampling sites discussed in the text. 1—Eocene–Oligocene lavas from the southern Qiangtang terrane; 2—late Eocene lavas and tuffs of the Linzizong Formation; 3—late Cretaceous redbeds of the Shexing Formation; 4—upper Cretaceous redbeds of the Takena Formation; 5—early Cretaceous lavas of the Woronggou Formation and Zenong Group. IYSZ—Indus-Yarlung suture zone. Lower left inset is shaded relief map of central Asia, showing the distribution of high elevation (darker shades) and political boundaries.

PALEOMAGNETIC DATA FROM THE LHASA TERRANE

Eocene–Oligocene Lavas

Paleomagnetic data from late Eocene–early Oligocene (35 \pm 3 Ma) basaltic lavas from the Qiangtang block were reported in Lippert et al. (2011). These lavas have been well dated using stepheating ⁴⁰Ar/³⁹Ar methods on bulk rock samples and hornblende or sanidine separates (Ding et al., 2007; Lippert et al., 2011). Site mean directions (n = 33) from three sampling localities have been filtered and binned into 20 statistically independent direction groups; the angular dispersion of the corresponding VGPs is consistent with modern secular variation studies. These lavas yield a maximum likelihood estimate of paleolatitude of $28.7^{\circ} \pm 3.7^{\circ}$ N at a present reference site at 33°N, 88°E on the Lhasa-Qiangtang suture. Because there was negligible north-south-directed upper crustal shortening within the Lhasa block throughout the Neogene (Kapp et al., 2005, 2007a, 2007b; Volkmer et al., 2007; Pullen et al., 2008; Rohrmann et al., 2012), we can transfer this paleolatitude directly to our reference position on the IYSZ, 4° of latitude to the south. This results in a paleolatitude estimate for the reference location on the southern margin of Asia of $24.7^{\circ} \pm 3.8^{\circ}$ N at 35 ± 3 Ma.

Paleogene Lavas: The Linzizong Formation

Basaltic to andesitic lavas, felsic tuffs, and volcaniclastic units of the Paleocene to early Eocene (64–47 Ma) Linzizong Formation blanket a large area of the southern Lhasa block (Fig. 1)

(He et al., 2007; Lee et al., 2009). The Linzizong Formation is the extrusive component of the Cenozoic part of the Gangdese magmatic arc (Mo et al., 2008), and it is variably divided into three (Dianzhong, Nianbo, Pana; Liu, 1993) or four (K-T, T1, T2, T3; He et al., 2007) subunits distinguished by petrology and unconformities. Although paleomagnetic data have been collected from all members of the Linzizong Formation (e.g., Chen et al., 2010), unfortunately too few published site mean directions from the lower and uppermost members meet our quality criteria to allow further consideration here (although, see Huang et al., 2013). Moreover, it is likely that sedimentary units below the T2 (He et al., 2007) and Pana Formation (Liu, 1993) have been remagnetized or have been affected by syndepositional and postdepositional inclination shallowing; the latter is almost certainly true for sedimentary units above the T2-Pana Formation (Huang et al., 2013). Here we only use paleomagnetic results from the felsic welded tuffs of the Pana unit (Liu, 1993) (unit T2 of He et al., 2007). This volcanic sequence is well dated as ca. 56-47 Ma using both wholerock ⁴⁰Ar/³⁹Ar (Lee et al., 2009) and zircon U/Pb (He et al., 2007) isotopic methods. Paleomagnetic data compiled in this study were collected from several separate localities of Linzizong Formation outcrops, but all are within ~200 km of the city of Lhasa (Fig. 1). Although Chinese geophysicists conducted paleomagnetic investigations of the Linzizong Formation from the early 1980s (e.g., Zhu et al., 1981), Achache et al. (1984) provided the first paleomagnetic results from the Linzizong Formation described in the English-language literature. Achache et al. (1984) sampled the Linzizong Formation along the Lhasa-Golmud highway between

TABLE 1. PALEOMAGNETIC POLES USED IN THIS STUDY Age N Age N ("N) ("E) ("A95 K S		cene-Ĕocene 51.5 ± 4.5 41 80.2 230.4 4.1 30.5 14.8 3.3 6.1 21.1 ± 4.1 b	r Cretaceous 86±14 <i>100</i> 74.6 346.5 2.7 − − 2.7 4.5 24.9 ± 2.7 c	r Cretaceous 97 ± 7 377 79.6 329.9 2.2 – – 1.6 2.0 23.7 ± 2.2 d	r Cretaceous 120 ± 10 13 16.2 ± 3.6 e, f	ian–Thanetian 59 ± 3 243 69.6 272.5 1.7 –	an-Maastrichtian 68 ± 3 144 55.8 261.6 3.5 2.3 3.6 −5.0 ± 3.5 g		0 24 88.5 173.9 1.9 29.1 ± 1.9 i	10 49 87.2 240.4 1.8 – – – 26.5 ± 1.8 i	20 31 83.7 254.7 2.6 – – – 22.9 ± 2.6 i	30 24 79.7 281.7 2.6 – – – 19.0 ± 2.6 i	40 24 74.7 286.8 2.9 – – – 14.4 ± 2.9 i	50 33 65.1 278.4 2.8 – – – 4.4 ± 2.8 i	60 44 48.5 280.8 2.1 – – – – – – 11.7 ± 2.1 i	70 32 36.4 280.7 2.5 – – – – – –23.5 ± 2.5 i	80 25 29.0 283.5 2.9 30.1 ± 2.9 i	90 28 20.9 291.4 2.5 35.2 ± 2.5 i	100 14 19.7 293.0 3.3 – – – – – – 35.7 ± 3.3 i	110 21 11.1 295.9 3.3 41.7 ± 3.3 i	120 28 8.6 296.4 2.6 43.5 ± 2.6 i	130 18 -1.0 297.1 2.850.6 ± 2.8 i		0 24 88.5 173.9 1.9 29.1 ± 1.9 i	10 49 86.7 150.0 1.8 30.5 ± 1.8 i	20 31 84.4 152.1 2.6 31.3 ± 2.6 i	30 24 83.1 146.5 2.6 − − − 32.4 ± 2.6 i	40 24 81.1 144.3 2.9 – – – – 33.6 ± 2.9 i	50 33 78.9 164.7 2.8 31.0 ± 2.8 i	60 44 78.2 172.6 2.1 – – – 29.4 ± 2.1 i	70 32 79.2 175.7 2.5 28.9 ± 2.5 i	80 25 79.7 177.9 2.9 28.5 ± 2.9 i	90 28 80.4 167.2 2.5 30.4 ± 2.5 i	100 14 80.8 152.3 3.3 32.6 ± 3.3 i	110 21 81.2 193.1 3.3 26.4 ± 3.3 i	120 28 79.0 190.1 2.6 26.2 ± 2.6 i	130 18 75.0 183.4 2.8 26.6 ± 2.8 i	.ocation—region from which paleomagnetic data were collected. N (N, n)—the number of paleomagnetic poles, virtual geomagnetic poles	n paleomagnetic poles listed in this table. A95—the radius of the 95% confidence cone on the mean pole; when calculated from samples	where do and un are the axes of the oval of 93% confidence about the fiteart pole calculated from the population of satriples. N—the t—the dispersion of the VGP non-lifetion A05 . / —the minimum and maximum values of A05 given N that the samnled non-lifetion of	Presented al 2011). A dash Phythem indicates argameter is not anniholded to this data set 1 + AI — nelected and GPS confidence		nole: calculations follow Butlar (1092). Sources: a	pole; calculations follow Butler (1992). Sources: a-Lippert et al. (2011); b-this study (see text and Table 2 for specific studies) with ages
TABLE 1. PALEOMAGNETIC Age N Pole N(N, n)	Eccene-Oligocene 35±3 20	⊃aleocene–Ĕocene 51.5 ± 4.5 41	Upper Cretaceous 86 ± 14 <i>100</i>	Upper Cretaceous 97 ± 7 377	Lower Cretaceous 120 ± 10 13	elandian–Thanetian 59 ± 3 <i>243</i>	npanian-Maastrichtian 68 ± 3 144		0 24	10 49	20 31	30 24	40 24	50 33	60 44	70 32	80 25	90 28	100 14	110 21	120 28	130 18		0 24	10 49	20 31	30 24	40 24	50 33	60 44	70 32	80 25	90 28	100 14	110 21	120 28	130 18	ath. Location-region from which paleomagnetic	mean paleomagnetic poles listed in this table. Ac	/ o.o., wriere <i>up</i> ariu <i>urri</i> are une axes or ure oval o ole S—the disnersion of the VGP nonulation Ao	iation (Deanen et al 2011) A dash (hvnhen) indi	from pole: calculations follow Butler (1992) Sour		
-igures 1 Location	Southern Lhasa block	Penbo	3 Maxiang	t Penbo	5 Deqing and Cuoqin	Tibetan Himalaya S	7 Zongshan Ca	Indian APWP															Asian APWP															Note: APWP-apparent polar wander	VGPs), or samples used to calculate the	auter utart poles of VGFS, A33 = (<i>up un</i> Tisher precision parameter of the mean	balenmarinetic data averades secular va	nterval for a reference site at 20°N 88°F		

Lhasa and Yangbajing, as well as near Linzhou, ~50 km northeast of Lhasa. Most subsequent paleomagnetic studies have concentrated on the exceptional exposures of the Linzizong Formation at Linzhou (Chen et al., 2010; Dupont-Nivet et al., 2010a; Liebke et al., 2010; Tan et al., 2010). A more regional picture is now available, however, with results from Mendui, ~25 km west-northwest of Linzhou (Sun et al., 2010), and from the Namling Basin near Shigatse, ~225 km southwest of Lhasa (Chen et al., 2010) (Fig. 1). The reported paleolatitude estimates for the reference location on the IYSZ from these individual studies range from ~8°N to as high as ~33°N, and these different estimates have been used to construct disparate tectonic models for the India-Asia collision and Tibetan Plateau. One of the goals of this study is to identify the source of this variability in these paleomagnetic data. We hypothesize that it is primarily the result of the low number (e.g., ~ 10) of independent site mean directions of high statistical quality acquired in each study.

Our compilation of reliable results includes five studies of the upper Linzizong Formation (i.e., T2–Pana Formation) listed in Table A1 in the GSA Data Repository¹ and shown in Figure 2. We have explicitly excluded other historical paleomagnetic studies of the Linzizong Formation (e.g., Pozzi et al., 1982; Westphal and Pozzi, 1983) because these reports have very small numbers of site mean directions, and because the more recent studies noted supersede all of these previous studies in terms of overall data quality and geomagnetic field representation. (For detailed rock magnetism and paleomagnetism of Linzizong Formation units sampled in each study, see the original publications; here we summarize the most salient observations.)

The magnetism of the T2-Pana member of the Linzizong Formation volcanic units is typically carried by magnetite to lowtitanium titanomagnetite with some high-titanium titanomagnetite or maghemite. These units often display two components of magnetization in orthogonal vector plots. A low-coercivity or low-temperature direction is indistinguishable from the modern field direction in in situ coordinates and is interpreted to record a viscous overprint. A higher coercivity or high-temperature direction that is distinct from the lower treatment level direction and shows univectoral decay toward the origin over several (>5) consecutive treatment steps on orthogonal vector plots defines the characteristic remanent magnetization (ChRM). In cases where maghemite has been detected, the low and high treatment level directions are not as distinct and the ChRM is determined using great circle direction fits (Dupont-Nivet et al., 2010a; Liebke et al., 2010). The tilt-corrected ChRM is often interpreted to be the primary remanent magnetization, although for most studies, the sampling strategy, results, or a combination of both preclude a rigorous paleomagnetic field test to clearly demonstrate this primary nature. Site mean directions from each study are presented in Table A1 (see footnote 1), and the mean VGPs and the corresponding latitude of the southern margin of the Lhasa block, at 29°N, 88°E, calculated from these poles are presented in Table 2.

There is considerable variation in not just the calculated paleolatitudes, as described here, but also the number of site mean directions used in each of the paleomagnetic data sets to calculate the mean poles (Table 2, Fig. 2). We suggest that the variability in the pole positions is a direct consequence of the small size of the individual sample sets, the inclusion of data that duplicate other spot readings, and the inclusion of spot readings that are not statistically equivalent (e.g., high versus low Fisher precision parameters), such that most of these individual studies do not provide a robust characterization of the time-averaged geomagnetic field. The dispersion (S) of VGPs calculated from the site mean directions for most of the individual studies is either greater or smaller than can be straightforwardly explained by the secular variation of the geomagnetic field (e.g., Johnson et al., 2008) and the A95values plot outside of the Deenen envelope (Deenen et al., 2011). These observations challenge the interpretation that these individual poles can be used for paleogeographic reconstructions.

We also note that Liebke et al. (2012) suggested that the mixed polarity observed in many of the dikes presented in Liebke et al. (2010) is primary and is the product of multiple injection episodes within each dike, such that each dike partially averages paleosecular variation. While we acknowledge this possibility, there is no reported independent field evidence for multiple injection episodes within the narrow (1-3 m) dikes. For simplicity, and in an effort to apply our selection criteria uniformly to all data sets, we will assume that each dike represents a single injection episode that can be treated equivalently to a lava unit. Notably, the exclusion of site mean directions from Liebke et al. (2010) from our final analysis does not significantly change the position of the mean paleomagnetic pole calculated for the Linzizong Formation and therefore does not change the paleolatitude of the southern margin of Asia ca. 50 Ma (see following).

We test the robustness of these individual paleomagnetic poles by combining site-mean directions from all of these studies and calculating a mean paleomagnetic pole from the VGP distribution (Fig. 2). If each individual study represents secular variation, then in principle, these individual poles should statistically overlap with the mean Linzizong Formation paleomagnetic pole. The data quality filter (described in Methods discussion) excludes 21 of 62 reported lava site mean directions (selecting from only sites used by the original authors of each study; Table A1 (see footnote 1). The mean paleomagnetic pole in tilt-corrected coordinates is located at $\lambda = 80.2^{\circ}$ N, $\phi = 230.4^{\circ}$ E (K = $30.5, A95 = 4.1^{\circ} (A95_{min} = 3.9^{\circ}; A95_{max} = 7.9^{\circ}), S = 14.8^{\circ}, n = 41)$ (Tables 1 and 2). The 41 site mean directions used to calculate this filtered paleomagnetic pole include data from each of the five independent studies and pass a regional fold test at 95% confidence (χ^2 test of McFadden, 1990). Although these data do not pass a reversal test at 95% confidence (class C: $\gamma_{calculated} = 15.8^{\circ}$, $\gamma_{\text{critical}} = 11.5^{\circ}$; McFadden and McElhinny, 1990), we note that

¹GSA Data Repository Item 2014216, paleomagnetic results from the Linzizong Formation (Table A1) and paleomagnetic results from Lower Cretaceous units from the Lhasa terrane (Table A2), is available at www.geosociety.org /pubs/ft2014.htm, or on request from editing@geosociety.org or Documents Secretary, GSA, P.O. Box 9140, Boulder, CO 80301-9140, USA.



Figure 2. Paleomagnetic data from the Linzizong Formation (LZF). Closed symbols indicate normal polarity directions; open symbols indicate reversed polarity directions. VGP—virtual geomagnetic pole. (A) Site mean directions from the LZF before the population has been filtered for consistent statistical quality (used by original authors of each study, N = 62). (B) Site mean directions from the LZF after the population has been filtered for consistent statistical quality (N = 41); in situ coordinates. (C) Same as B, except in tilt-corrected coordinates. (D) Paleomagnetic pole for the LZF calculated from the virtual geomagnetic poles (VGPs) of the paleomagnetic data shown in C. In the legend, *n* refers to the number of spot readings provided by each study; number in parentheses is number of spot readings that meet our quality criteria.

the reversal test may not be ideally suited for sequences of extrusive rocks because individual eruptive events and even subsets of eruptive events (e.g., lavas that erupted during a single reversed or normal polarity interval) may not provide a time-averaged representation of the geomagnetic field (i.e., average paleosecular variation). The fact that a normal-reverse-normal polarity sequence is recorded within a long interval of lavas, however, is strong evidence that the lavas record primary directions and record several hundreds of thousands to millions of years of time. It is important that the calculated dispersion of this filtered mean VGP is consistent with known dispersion caused by paleosecular variation (Johnson et al., 2008), the A95 is within the confidence envelope of Deenen et al. (2011), and the pole fulfills all seven quality criteria of van der Voo (1990). We conclude that this paleopole more accurately characterizes the time-averaged behavior of the geomagnetic field than any previous individual study of the Linzizong Formation.

Notably, only in Dupont-Nivet et al. (2010a) were paleomagnetic data assessed for direction grouping. Ideally, we would like to assess all of the Linzizong Formation site mean directions for appropriate direction grouping to ensure that only independent samples of the geomagnetic field direction are averaged, but we are currently limited by the stratigraphic fingerprinting of regionally extensive lavas in the Linzizong Formation. For example, Tan et al. (2010) presented nine site mean directions over a short interval in outcrop, but only six of these directions are

			Pole latitude	Pole longitude								
Study	Filtered	Ν	(°N)	(°E)	A95	K	S	$A95_{min}$	$A95_{max}$	I _p	±	ΔI_p
Achache et al. (1984)	no	7	72.7	287.6	8.8	47.7	11.8	7.8	24.1	12.6	±	8.8
	yes	4	72.5	283.4	8.6	114.6	7.6	9.8	34.2	12.1	±	8.6
Tan et al. (2010)	no	9	89.6	271.6	6.6	62.0	10.3	7.1	20.5	28.6	±	6.6
	yes	9	89.6	271.6	6.6	62.0	10.3	7.1	20.5	28.6	±	6.6
Dupont-Nivet et al. (2010a)	no	31	76.7	208.6	4.7	31.3	14.5	4.3	9.4	21.7	±	4.7
	yes	24	77.6	211.6	5.0	35.7	13.6	4.8	11.1	21.7	±	5.0
Liebke et al. (2010)	no	10	70.2	210.6	12.9	17.0	20.1	6.8	19.2	17.2	±	12.9
	yes	1	68.1	224.2	_	_	_	_	-	12.4	±	_
Sun et al. (2010)	no	14	74.0	274.4	8.4	23.1	16.9	5.9	15.6	13.1	±	8.4
	yes	3	67.7	268.2	23.9	27.6	15.5	11.0	41.0	6.7	±	23.9
Mean pole (this study)	no	62	79.1	236.8	3.8	23.1	17.1	3.3	6.1	19.5	±	3.8
	yes	41	80.2	230.4	4.1	30.5	14.8	3.8	7.9	21.1	±	4.1
	yes*	17	80.9	269.5	6.4	32.3	14.3	5.5	13.8	19.9	±	6.4

Note: Conventions and symbols as in Table 1. Filtered: mean virtual geomagnetic pole (VGP) calculated without or with applying the data quality filter described in the text (yes*—filtered pole without the results of Dupont-Nivet et al., 2010a). Dashes (blanks) in Liebke et al. (2010) column are because this pole results from only one VGP.

stratigraphically independent (following Mankinen et al., 1985). Moreover, the results from Tan et al. (2010) overlap with a few of the stratigraphically highest sites reported in Dupont-Nivet et al. (2010a). Thus, our new pole, which incorporates all statistically robust site mean data from paleomagnetic studies of the Linzizong Formation, almost certainly contains some non-independent data, but we maintain that this remains the most comprehensive and statistically consistent pole for the Linzizong Formation. The corresponding 56–47 Ma paleolatitude of the reference site on the IYSZ is $21^{\circ} \pm 4^{\circ}$ N.

Inspection of Figure 2 and Tables 2 and A1 (see footnote 1) suggests that the filtered data set favors our own results (Dupont-Nivet et al., 2010a). We note that this is the result of the large size of the Dupont-Nivet et al. (2010a) data set and that the same quality control in the original study was applied in this study, so no direction groups were culled from the compilation. To test if the large data set of Dupont-Nivet et al. (2010a) biases the mean Linzizong Formation paleopole, we recalculated the pole using only the filtered data described here and excluding the site mean directions from Dupont-Nivet et al. (2010a). The resulting pole is calculated from 17 tilt-corrected site mean directions and is located at $\lambda = 80.9^{\circ}$ N, $\phi = 269.5^{\circ}$ E (K = 32.3, A95= 6.4°; A95_{min} = 5.5° ; A95_{max} = 13.8°, S = 14.3°). A simple F-test indicates that the filtered Linzizong Formation poles with and without the Dupont-Nivet et al. (2010a) data are not significantly different at high confidence levels (p = 11.48%). We conclude that the large data set of Dupont-Nivet et al. (2010a) does not significantly bias the mean Linzizong Formation pole.

In summary, we have combined site mean directions provided by six independent paleomagnetic studies of the Linzizong Formation to calculate a new paleomagnetic pole for the Lhasa terrane in early Eocene time. After filtering the original 62 site mean directions to produce a data set composed of statistically consistent site mean directions, we calculated a mean pole located at $\lambda = 80.2^{\circ}$ N, $\phi = 230.4^{\circ}$ E (K = 30.5, A95= 4.1°, S = 14.8°) (Fig. 2). The 41 underlying site mean directions used to calculate this pole pass a paleomagnetic fold field test at high confidence, exhibit a reversal stratigraphy, are characterized by a VGP dispersion that is in excellent agreement to that observed in 0-5 Ma lavas sampled at the same latitude at Lhasa, and exhibit scatter in the data set that can be straightforwardly explained by paleosecular variation. We conclude that this pole is the most statistically comprehensive and coherent paleomagnetic pole from Tibetan Asia for the Cenozoic and is ideally suited to provide (1) the paleolatitude of the Tibetan Plateau, (2) precise paleomagnetic constraints on the age and latitude of the collision between the Tibetan Himalaya and the Lhasa terrane, and (3) the distribution of convergence within the India-Asia collisional orogen.

Late Cretaceous Redbeds, Maxiang

Sun et al. (2012) reported paleomagnetic data from the $86 \pm$ 14 Ma Shexing Formation of redbeds that crop out along the Lhasa-Golmud highway near Maxiang (Fig. 1). Site mean direc-

tions calculated from the high-temperature ChRM component pass a fold test at high confidence. We evaluated the distribution of high-temperature specimen directions for inclination shallowing following the elongation/inclination (E/I) method of Tauxe and Kent (2004) (original ChRM directions provided by Zhiming Sun, 2011, personal commun.). Samples with mean angular deviations of ChRM directions $>15^{\circ}$ (n = 5) and with reversed polarity ChRM (n = 2) directions were removed from the population to exclude poorly resolved or transitional directions. The Vandamme cut-off procedure for identifying outliers in the direction population (Vandamme, 1994) removed another seven samples, resulting in a mean VGP for the Shexing Formation located at $\lambda =$ $70.5^{\circ}N, \phi = 324.9^{\circ}E$ (samples corrected = 100, K = 32.5, A95 = 2.5° ; A95_{min} = 2.7° ; A95_{max} = 4.5° , S = 14.3°). We calculated a flattening factor of 0.66 (0.53 and 0.91 95% confidence interval; Fig. 3), resulting in an E/I corrected mean VGP located at $\lambda =$ 74.6°N, $\phi = 346.5$ °E (N = 100, K = 28.6, A95= 2.7°; A95_{min} = 2.7°; A95_{max} = 4.5°, S = 14.3°). These results suggest ~9° of inclination shallowing in the fluvial-lacustrine redbeds of the Shexing Formation, which is typical of these facies in Asia (Dupont-Nivet et al., 2010b; Lippert et al., 2011). The calculated A95 is at the lower limit of the minimum value predicted by geomagnetic field models; i.e., paleosecular variation is at the threshold of being represented by this data set. This can be straightforwardly explained by some averaging of secular variation within individual specimens, which is common in sediments and especially redbeds (Deenen et al., 2011). The calculated paleolatitude for the reference location on the IYSZ is $24.9 \pm 2.7^{\circ}$ N at 86 ± 14 Ma.

Late Cretaceous Redbeds, Penbo

We use paleomagnetic data from the nonmarine Takena Formation and calculations of a Late Cretaceous (97 \pm 7 Ma; Leier et al., 2007a) paleolatitude for the Lhasa block of Tan et al. (2010). The distribution of the sample directions from this data set is consistent with inclination shallowing that is readily attributed to sedimentary depositional processes. Tan et al. (2010) corrected these directions using the E/I method. The paleomagnetic pole calculated from these tilt and inclination shallowing corrected sediments is located at 79.6°N, 329.9°E (N = 377, A95= 2.2; A95_{min} = 1.6° ; A95_{max} = 2.0°) (Table 1). The calculated A95 is slightly higher than expected from paleosecular variation alone, suggesting some additional error due to local rotations between sampling localities, bedding measurement errors, the inclusion of outlier directions, or a combination of these factors. Taking these data at face value, the paleolatitude of the reference location on the IYSZ calculated from this pole is $23.7^{\circ} \pm 2.2^{\circ}$ N ca. 97 ± 7 Ma.

Tan et al. (2010) reported paleomagnetic data from what they described as 32 consecutive, 2–30-cm-thick lavas interbedded with the Takena Formation. We do not use these data because 2–30 cm is unusually thin for lavas and no pictures or further descriptions of these rocks were provided, and because our own field observations, as well as the comprehensive field study of



Figure 3. Elongation/inclination (E/I) correction (Tauxe and Kent, 2004) of high-temperature, primary magnetization component of the Shexing Formation redbeds at Maxiang, south-central Lhasa terrane (Sun et al., 2012). D—declination; n—number of samples; α_{o5} —half angle of the 95% confidence cone on the mean direction. (A) Equal area plot of original paleomagnetic directions in tilt-corrected coordinates, with a mean inclination (I_o) of 31.2° (gray star, 95% confidence interval indicated in surrounding light gray). Slight horizontal elongation of the direction distribution is consistent with sedimentary inclination shallowing. (B) E/I-corrected paleomagnetic directions after accounting for a flattening factor of 0.66. The E/I-corrected inclination (I_o) is 41.6°. (C) Plot of elongation versus inclination for the TK03.GAD (Tauxe and Kent, 2004) paleosecular variation model (bold dashed gray line) and for the Shexing Formation directions (bold barbed black line) for different flattening factors, f. Thin light gray lines are the results from 20 boot-strapped data sets. The crossing points of the Shexing Formation and bootstrapped data set with the TK03.GAD results represents the elongation/inclination pairs that are most consistent with the TK03.GAD field model. (D) Histogram of crossing points from 5000 boot-strapped data sets. The most frequent flattening factor is f = 0.66, resulting in the distribution of directions shown in B.

Leier et al. (2007a), did not identify any lavas within the Takena Formation. There are, however, widespread sills and dikes (e.g., He et al., 2007; Liebke et al., 2010) commonly with early Eocene 40 Ar/ 39 Ar whole-rock (Yue and Ding, 2006) and U/Pb zircon ages (He et al., 2007).

We also note that we have not used paleomagnetic data from the Takena Formation reported by Lin and Watts (1988). The reasons for this are (1) Lin and Watts (1988) provided only site mean directions from their sampling sites, whereas individual sample ChRM directions are needed to evaluate these results for sedimentary inclination shallowing. (2) Tan et al. (2010) resampled some of the same horizons as Lin and Watts (1988). Inclusion of these results would lead to oversampling of the geomagnetic field, resulting in spurious or misleading sampling statistics. It is neither necessary nor justified to include the Takena paleomagnetic from Lin and Watts (1988) in our compilation.

Early Cretaceous Lavas

Paleomagnetic results from volcanic rocks of Cretaceous age have been reported in two recent studies. Sun et al. (2012) reported 15 site mean directions from the 114.2 ± 1.1 Ma Woronggou Formation rhyolites near Deqing, on the west shore of Nam Co (Fig. 1; Table A2 [see footnote 1]); only five site mean directions meet our data quality standards. Chen et al. (2012) reported 19 site mean directions from 120 ± 10 Ma tuffs and lavas of the Zenong Group near the town of Cuoqin in the central Lhasa terrane (Fig. 1; Table A2 [see footnote 1]); only nine site mean directions describe the geomagnetic field at this locality after the data are filtered for consistent quality.

We combine the filtered results of these two studies to calculate a lava-based paleomagnetic pole from the Early Cretaceous of the Lhasa terrane. Although both localities show similar inclinations, the declinations vary by ~ 30° – 40° , implying significant rotations between the two sampling regions (Fig. 4). We follow Doubrovine and Tarduno (2008) and calculate a maximum likelihood virtual geomagnetic latitude (VGL) using the Arason and Levi (2010) method. First, we calculate the VGPs for the filtered site mean directions and then calculate the VGL of the reference site on the IYSZ. These VGLs are then used as input to calculate the maximum likelihood latitude for the reference site from these data: $16.2^{\circ} \pm 3.6^{\circ}$ N at 120 ± 10 Ma. We note, however, that the dispersion of these VGLs is only 7.0°, suggesting that paleosecular variation has been underrepresented by the lavas. We caution that this paleolatitude estimate maybe not be robust.

IMPLICATIONS: DISCUSSION

Latitude of the Southern Lhasa Terrane since the Early Cretaceous

The paleomagnetic data described above record the latitude of the southern margin of the Lhasa terrane since ca. 110 Ma (Fig. 5). These latitudes are calculated from paleomagnetic data from well-dated lavas and sediments that mostly appear to average short-term geomagnetic field behavior (i.e., secular variation) and have been corrected for sedimentary inclination shallowing biases. We emphasize that the paleolatitude curve shown in Figure 5 is calculated from the statistically most comprehensive and consistent paleomagnetic data set currently available for the Lhasa terrane from 85° to $91^{\circ}E$.

Figure 5 shows that the Lhasa terrane has remained relatively stable in terms of latitude since the Early Cretaceous. Notably, none of the paleomagnetic data sets is consistent with low, tropical latitudes for the Lhasa terrane, in contrast to conclusions drawn from some individual studies based on small data sets associated with high dispersion (Westphal et al., 1983; Lin and Watts, 1988; Chen et al., 2010). Instead, the Lhasa terrane was positioned at high tropical to low subtropical north latitudes (16°-22°N) since ca. 110 Ma and has translated northward to its present latitude at 29°N only in the past 50 m.y. Part of this motion can be attributed to the absolute plate motion of Eurasia relative to the spin axis described in the GAPWP (Torsvik et al., 2012), and part is related to north-south shortening of the Asian plate, e.g., in the Tibetan Plateau and the Tien Shan (e.g., Molnar and Tapponnier, 1975; Dewey et al., 1989; van Hinsbergen et al., 2011a). This latitude history is a key boundary condition for interpreting the kinematics and geodynamics of the India-Asia convergence history and tectonics and climate of the Lhasa terrane and surrounding region. Each of these topics is discussed in the following.

Paleomagnetic Estimate of the Age of the Tibetan Himalaya–Lhasa Collision

The onset of continental collision between India and Asia is of paramount interest to many Earth scientists because it is conventionally thought to mark the onset of surface uplift of the



Figure 4. Equal area plot of site mean directions from Early Cretaceous volcanic rocks from the central and western Lhasa terrane that pass our data quality criteria (Sun et al., 2008; Chen et al., 2012). (A) In situ coordinates. (B) Tilt-corrected coordinates. Downward pointing directions are shown as filled symbols; upward pointing directions are shown as open symbols. Significant improvement in the clustering of inclinations from in situ coordinates to tilt-corrected coordinates is consistent with a prefolding magnetization. The noticeably different declination values observed in the two localities are consistent with differential vertical axis rotation since the emplacement of these lavas and justify the use of latitude-only data analysis methods (see text for details).

Tibetan Plateau (but see following discussion), and because it is the archetypal setting for many models about the dynamics of continental subduction (e.g., Chemenda et al., 2000; Capitanio et al., 2010). Moreover, the demise of the Neotethyan Ocean should impart significant changes in plate force balances and ocean circulation. The age of the onset of collision between continental units of India and Asia is a key boundary condition for understanding the geodynamics governing the distribution of strain in subduction zones and collisional orogens, because, with the use of the plate circuits (e.g., Molnar and Stock, 2009; Copley et al., 2010; van Hinsbergen et al., 2011b), it sets the budget for the magnitude of plate convergence that must be accounted for within the orogen. It is this last point that makes knowing the India-Asia collision age especially important.

The end of marine sedimentation within the Tibetan Himalaya (Rowley, 1996; Garzanti, 1999; Green et al., 2008; Najman et al., 2010; Hu et al., 2012) in the early Eocene and the appearance of Asian detritus in Tibetan Himalaya sediments by 52 Ma (Najman et al., 2010; Wang et al., 2011; Hu et al., 2012) are generally considered to reflect the accretion of the Tibetan Himalaya to the Asian plate at that time (see review by Najman et al., 2010). This time period also corresponds to a nearly twofold decrease in the relative motion rates between the Indian and Asian plates, which has been interpreted to represent the arrival of buoyant, extended continental margin crust to the subduction zone beneath southern Tibet, as well as increased resistance at the subduction zone due to surface uplift above the trench (Iaffaldano and Bunge, 2009; van Hinsbergen et al., 2011b).

Paleomagnetism can be used to independently estimate the Tibetan Himalaya-Asian collision age by determining when paleolatitudes of the Tibetan Himalaya and southern Tibet start to overlap. Previous paleomagnetic-based arguments have produced Tibetan Himalaya-Asia collision ages that range from 70 Ma (Klootwijk et al., 1992) to ca. 35 Ma (Ali and Aitchison, 2006) and several ages between (e.g., see individual Linzizong Formation studies noted above). The disparity of collision ages results from (1) the use of data from sedimentary rocks that have not been evaluated for sedimentary inclination shallowing; (2) the use of data from lavas that insufficiently represent secular variation; or (3) using APWPs of the stable continents, or only selected data from India and Eurasia while disregarding plate circuits altogether and inferring dimensions of Greater India and greater Asia without taking paleomagnetic evidence into account (e.g., Ali and Aitchison, 2006).

The paleolatitude history of the Lhasa terrane described here and a recent reappraisal of high-quality paleomagnetic data of primary origin from Late Cretaceous to late Paleocene marine sediments from the southern Tibetan Himalaya (Patzelt et al., 1996; Yi et al., 2011; see also van Hinsbergen et al., 2012b) can be used to estimate a paleomagnetic collision age. Using a reference point located at 29°N, 88°E on the IYSZ, the latitude of the northern Tibetan Himalaya overlaps with the latitude of the southern margin of the Lhasa terrane at 49.4 ± 4.5 Ma, at lat 21° ± 4°N (95% confidence intervals; see Fig. 5A). This age is Figure 5 (*on following page*). The age and latitude of the Tibetan Himalaya–southern Tibet collision as calculated from paleomagnetic data discussed in the text and presented in Table 1. Paleolatitude versus time is shown for a point on the modern Indus-Yarlung suture zone (29°N, 88°E) in Eurasia (gray), southern Lhasa terrane, Tibetan Himalayan, and India reference frames. APWP—apparent polar wander path. (A) Without accounting for crustal shortening within the Indus-Yarlung suture zone (IYSZ) during and since collision, paleomagnetic data indicate that the collision between the Tibetan Himalaya and southern Tibet occurred ca. 49.5 ± 4.5 Ma at $21^{\circ} \pm 4^{\circ}N$ (95% confidence interval). (B) With conservative estimates of the magnitude of crustal shortening (see text for details), the paleomagnetically determined collision age is ca. 52.5 ± 4.5 Ma at $19^{\circ} \pm 4^{\circ}N$ (95% confidence interval).

entirely consistent with estimates derived from geologic observations for the collision described.

We emphasize that this paleomagnetically determined collision age is a minimum estimate. The paleolatitudes calculated here describe the location of the suture that currently separates the Tibetan Himalaya from Asian crustal units. It is likely, however, that prior to collision, the Neotethyan subduction zone was separated from Asia by a forearc. At present, the southern Lhasa terrane immediately adjacent to the suture hosts the Gangdese volcanic arc (of which the Linzizong Formation is part), which formed since Early Cretaceous time during northward subduction of the Neotethyan Ocean (Ji et al., 2009). Prior to collision, however, a forearc must have separated the Gangdese arc from the subduction zone; relics of a Cretaceous forearc are preserved in the IYSZ (Einsele et al., 1994; Yin et al., 1994). There are no direct constraints on the original width of this forearc, but typical arc-trench distances today are ~200-300 km wide (modern observations using GeoMapApp; http://www.geomapapp.org). The initial collision between the Tibetan Himalaya and southern Tibet may have occurred $\sim 2^{\circ}$ south of the modern southern boundary of the Lhasa terrane (i.e., the suture). After accounting for the shortened Gangdese forearc, we predict that the southern margin of Asia may have been located at $19^\circ \pm 4^\circ N$ ca. 50 Ma.

In addition, the collision age is calculated based on the modern northernmost exposures of Paleocene rocks of the Tibetan Himalaya. If Tibetan Himalayan rocks existed farther north, but were underthrusted entirely below the Lhasa terrane, initial collision could have been older. For example, crustal shortening of the Tibetan Himalayan sequence north of Duela and Gamba, where the Tibetan Himalaya paleomagnetic data were collected, is unconstrained by detailed field studies; it may be as much as 50% or more (Ratschbacher et al., 1994; Ding et al., 2005), accounting for another ~1° of more northern latitude for the leading edge of the Tibetan Himalaya terrane prior to collision.

We conservatively add 2° of latitude to the leading edge of both the Tibetan Himalaya and southern Tibet to estimate the effect of postcollisional shortening within the IYSZ on the paleomagnetic collision age. Given the very high plate convergence rates between 60 and 50 Ma (16–18 cm/yr; van Hinsbergen et al., 2011b), these modifications would have only a minor effect on

Early Cretaceous-present latitude of the proto-Tibetan Plateau



the collision age, which would be pushed back to 52.4 ± 4.5 Ma at $19^{\circ} \pm 4^{\circ}N$ (95% confidence). We emphasize that although these shortening corrections appear reasonable, the actual magnitude of shortening within the Xigaze forearc, IYSZ, and Tibetan Himalaya fold-thrust belt remains poorly determined by geologic studies. Nevertheless, the mean IYSZ shortening-corrected collision age is in even better agreement with the stratigraphic record of collision, which indicates that the Tibetan Himalaya was close enough to Asia by 52 Ma to be a sink for Asian-derived detritus (e.g., Najman et al., 2010; Wang et al., 2011; Hu et al., 2012). The preponderance of geological and geophysical data from the IYSZ indicates that collision between continental units of Indian affinity with the southern margin of Asia was ongoing by 52 Ma. Until improved estimates of crustal shortening within the suture zone become available, we take a conservative approach and conclude that this collision occurred at $\sim 21^{\circ} \pm 4^{\circ}$ N.

Distribution of Strain in the India-Asia Collisional Orogen

Collision at 52 Ma implies that ~2860-3600 km of India-Asia plate convergence (from the western to eastern syntaxes) must be accounted for in the Himalaya-Tibetan orogen (van Hinsbergen et al., 2011b, 2012b). However, the plate circuit from which these convergence budgets are calculated cannot predict where this convergence is distributed, and therefore how and where strain accumulates in the India-Asia collisional orogen. The paleolatitude record for the Lhasa terrane described herein provides a fundamental, first-order constraint on this problem by locating the southern margin of Asia within the India-Asia plate circuit. Comparing the ca. 52 Ma collision of lat $\sim 21^{\circ} \pm$ 4°N calculated for the IYSZ in southern Tibet coordinates to the $31^{\circ} \pm 3^{\circ}$ N paleolatitude predicted from the GAPWP in Eurasian coordinates (Torsvik et al., 2012) suggests that $\sim 1100 \pm 560$ km of convergence was consumed between southern Tibet and stable Asia. This magnitude of convergence is larger than the amount of shortening (600-1500 km from eastern Tibet to the Pamir) north of the IYSZ since 50 Ma, calculated from structural geologic estimates (van Hinsbergen et al., 2011a). However, it has long been recognized that paleomagnetic inclinations across Central and East Asia (e.g., in Mongolia, north of the region that accommodated the majority of Cenozoic shortening related to India-Asia convergence; Hankard et al., 2007; Dupont-Nivet et al., 2010b) are also significantly shallower in Cenozoic time than predicted by the GAPWP. Therefore, the calculated paleolatitude difference computed from the paleolatitude of the Lhasa terrane compiled here, and that predicted by the GAPWP (Torsvik et al., 2012), cannot entirely be ascribed to India-Asia collision-related shortening of Tibet. Previous proposals to explain the Asian inclination shallowing include (1) non-dipole contributions (Si and Van der Voo, 2001), (2) mobility of Asia relative to Eurasia along cryptic faults in the Urals and along the Tornqvist-Tesseyre line (Cogné et al., 1999, 2013; Hankard et al., 2007), or (3) a combination of time-dependent non-dipole contributions and imperfections in the GAPWP (Dupont-Nivet et al., 2010b). For a detailed discussion of the Asian inclination shallowing problem, including possible solutions and important implications, there are reviews (Dupont-Nivet et al., 2010b; Lippert et al., 2011). What is clear from the paleomagnetic and structural studies (Dewey et al., 1989; Searle et al., 2011; van Hinsbergen et al., 2011a) of the Tibetan orogen is that the magnitude of convergence and upper crustal shortening within Asia since ca. 50 Ma is on the order of hundreds, not thousands of kilometers, as once assumed based on pioneering paleomagnetic and structural studies (e.g., Achache et al., 1984; Lin and Watts, 1988; Tapponnier et al., 1982).

A direct, but often overlooked, implication from nearly all studies of the Tibetan orogen is that all convergence since 50 Ma that was not accommodated north of the IYSZ by intra-Asian deformation must consequently have been accommodated south of the IYSZ by Indian plate subduction. The paleomagnetic data from the Lhasa terrane reviewed herein, and those data compiled from the Tibetan Himalaya (Patzelt et al., 1996; Yi et al., 2011; see also van Hinsbergen et al., 2012b) show that as much as \sim 2440 ± 330 km of convergence was accommodated between the Tibetan Himalaya and Himalayan front. This amount of convergence far exceeds the minimum shortening estimates for the Greater and Lesser Himalayan thrust belts (400-700 km; DeCelles et al., 2002; Long et al., 2011), and those of the few structural studies of the Tibetan Himalaya (100-200 km; Ratschbacher et al., 1994; Wiesmayr and Grasemann, 2002; Murphy and Yin, 2003). It follows from the paleomagnetic reconstructions that a large part of Greater India lithosphere has no as-yet recognized record at the surface (due largely to a combination of erosion and underthrusting). Paleomagnetic and plate kinematic data clearly indicate that in Late Cretaceous time Greater India was considerably more extensive (as much as 3000 km beyond modern cratonic India) than in the Early Cretaceous and before (<900 km) (van Hinsbergen et al., 2012b). Van Hinsbergen et al. (2012b) argued that Cretaceous Greater India consisted of thinned continental and oceanic lithosphere (Greater India Basin) resulting from a ca. 120–70 Ma rifting event that separated a microcontinent carrying the Tibetan Himalayan rocks from the Indian craton. Others have previously suggested extension within Greater India, although the timing and the implications for the India-Asia collision are different (Roy, 1976; Appel and Li, 1988; Hsu et al., 1995; Patzelt et al., 1996). Van Hinsbergen et al. (2012b) concluded that the ca. 52 Ma Tibetan Himalaya-Asia collision was characterized by a microcontinent-continent collision (Fig. 6), followed by subduction of the Cretaceous Greater India Basin beneath the southern margin of Asia, and then by collision between cratonic India and Asia at 25-20 Ma.

Surface Uplift of the Tibetan Plateau, Paleogeography of the India-Asia Collision Zone, and Implications for Atmospheric Dynamics and Regional Climate

The onset of surface uplift of the Tibetan Plateau from near sea level is generally considered to have begun shortly after the onset of collision between the Tibetan Himalaya and southern



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Figure 6. Early Eocene paleogeographic reconstruction of the Neotethyan region at the time of collision between the Tibetan Himalayan microcontinent and the southern margin of Asia. Key distinctions between this reconstruction and previous reconstructions include the elevated Lhasaplano (Kapp et al., 2005) (i.e., proto–Tibetan Plateau) and the oceanic Greater India Basin (van Hinsbergen et al., 2012b). Neotethyan subduction zone (bold red line) reconstruction is after van Hinsbergen and Schmid (2012), van Hinsbergen et al. (2012b, 2014), Gaina et al. (2013), Agard et al. (2011), Argnani (2012), and Hall (2012). Paratethys region (e.g., Turan Sea, Tarim Basin) reconstruction is after Bosboom et al. (2011, and references therein).

Tibet, such that the Tibetan Plateau is often considered to be the result (almost) entirely of the India-Asia collision (Molnar and Tapponnier, 1975; England and Houseman, 1986; Tapponnier et al., 2001; Royden et al., 2008; Clark, 2012). However, several lines of geologic evidence challenge this simplified view of the surface uplift history of the Tibetan Plateau. Northward subduction of the Neotethys Ocean below the Lhasa terrane began shortly after the collision of the Lhasa terrane with the Qiangtang terrane to the north in latest Jurassic–earliest Cretaceous time (Maluski et al., 1982; Leier et al., 2007b; Zhu et al., 2009). Also in Cretaceous time, and so well before the Tibetan Himalaya–Lhasa collision, well-documented fold and thrust fault belts developed on the Lhasa terrane (~250 km of north-south shorten-

ing) and Qiangtang terrane (~400 km of north-south shortening) (Kapp et al., 2005, 2007a, 2007b; Leier et al., 2007a; Volkmer et al., 2007; Pullen et al., 2008). Marine sedimentation throughout the northern Lhasa and southern Qiangtang terranes ceased during Aptian time (Zhang, 2000; Kapp et al., 2007a; Leier et al., 2007a), even though this period of time corresponds to global sea-level highstands (Müller et al., 2008), indicating that crustal thickening had been sufficient to raise the surface elevation of Tibet above sea level since ca. 110 Ma. Thus, a proto–Tibetan Plateau probably developed in Cretaceous-Paleogene time and characterized the physiography of the southern margin of Asia from at least 45 Ma (Dupont-Nivet et al., 2008; Wang et al., 2008; Rohrmann et al., 2012), and probably even Late Cretaceous time

(England and Searle, 1986; Murphy et al., 1997; Kapp et al., 2005, 2007a, 2007b; Searle et al., 2011). The stable oxygen isotope paleoaltimetry record from Tibet, albeit limited and reliant on numerous assumptions, is generally consistent with an Eocene proto-plateau (see recent review by Quade et al., 2011). Therefore, a significant part of the shortening within (van Hinsbergen et al., 2011a, and references therein) and surface uplift (see references in preceding two sentences) of the Tibetan Plateau predate the Tibetan Himalaya-Lhasa collision. We note that an important implication of the two-stage India-Asia collision as proposed by van Hinsbergen et al. (2012b) is that almost all surface uplift and shortening of Tibetan lithosphere, except for northeastern and possibly eastern Tibet (e.g., Kirby et al., 2002; Dettman et al., 2003; Hough et al., 2011; Zhuang et al., 2011; Lease et al., 2012) probably predate the final continent-continent collision (ca. 25-20 Ma). Numerous geologic studies indicate that crustal shortening and surface uplift in northeastern and eastern Tibet were also substantial before this time (Jolivet et al., 2001; Horton et al., 2002, 2004; Richardson et al., 2008; Yan et al., 2011; Zhuang et al., 2011; Wang et al., 2012). Although the shortening and surface uplift of the Tibetan Plateau is usually ascribed to the collision between India and Asia, the evidence and arguments listed here instead suggest that the Tibetan Plateau may be quite similar to the Altiplano of the Andes, in that most crustal shortening and plateau uplift developed above an oceanic subduction zone in the absence of a continent-continent collision. Also similar to the Andes, most of the proto- and modern Tibetan Plateau developed at dry, subtropical latitudes.

Another important implication of the paleolatitude of the Lhasa-plano (i.e., proto-Tibetan Plateau) during the two-stage India-Asia collision history described in van Hinsbergen et al. (2012b) is that throughout the Paleogene a deep tropical to subtropical oceanic basin remained between the high topography of the proto-Himalaya and Lhasa-plano to the north and continental India to the south. This Greater India Basin (van Hinsbergen et al., 2012b) would have connected the Indonesian seaways to the eastern Mediterranean Tethys (Fig. 6). Thus, east-west oceanic and atmospheric circulation in the Tethys may have been sustained much longer than normally assumed, although the details of through-going circulation depend on the latitude and depth of the seaway. Both circulation at tropical to subtropical latitudes and the presence of a deep and large body of water undoubtedly had a prolonged and profound effect on the paleoclimate in the region, particularly in terms of the Indian monsoon, physical and chemical weathering, and carbon burial.

For example, the modern plateau exerts a strong control on the distribution of water vapor (and therefore heat) on the planet by diverting and at times focusing or blocking major atmospheric jets such as the ITCZ and Westerlies (Wallace and Hobbs, 2006; Molnar et al., 2010). Jet diversion around the margins of the plateau affects not only the regional climate, but also areas far downstream, such as in the eastern Mediterranean and Africa (Hoskins and Karoly, 1981; Rodwell and Hoskins, 1996). How sensitive is the climate system to the latitudinal position of the Tibetan

orographic barrier and the distribution of proximal tropical water masses upstream of the jets? Does the location of the steep orographic barrier that may be required to initiate the Indian monsoon (Boos and Kuang, 2010) modulate the seasonal diversion of the ITCZ, intensity of the monsoon, and its teleconnections downstream (Hoskins and Karoly, 1981; Wallace and Hobbs, 2006; Huber and Goldner, 2012)? Did the presence of the Greater India Basin, and therefore tropical throughflow from the Pacific to the subtropical paleo-Arabian Sea and eastern Mediterranean Tethys, enhance the distribution of seasonal precipitation and amplify continental weathering along the margins of these water masses (e.g., Kent and Muttoni, 2008, 2013; Bookhagen et al., 2005)? If there is a strong link between the paleolatitude of the Lhasa-plano, the Greater India Basin, and precipitation patterns in and around Tibet, then we should expect feedbacks with vegetation and erosion, which will in turn alter ocean chemistry (Misra and Froelich, 2012; Pälike et al., 2012), and perhaps even the carbon cycle (France-Lanord and Derry, 1997; Kent and Muttoni, 2008, 2013; Allen and Armstrong, 2012) if physical and chemical erosion rates are high. Prolonged and enhanced seasonal precipitation in Eocene and Oligocene time, fueled in part by the steep and high topographic barrier of the proto-Himalaya and Tibetan Plateau, but also by the presence of the tropical Greater India Basin (and the presence of the para-Tethys in Central Asia (Ramstein et al., 1997; Bosboom et al., 2011), is also expected to have substantially modified the distribution and isotopic composition of precipitation on the proto-Himalaya foreland and Tibetan Plateau. These environmental factors must be considered when interpreting archives of ancient precipitation to infer the paleoaltimetry of Tibet in deep time. From a global perspective, future climate sensitivity studies should explore the direct and secondary effects of a larger, less circuitous tropical seaway on Cenozoic climate change. The possible relationships between the demise of the Greater India Basin, the carbon cycle, and Eocene cooling may be particularly important (Hotinski and Toggweiler, 2003; von der Heydt and Dijkstra, 2008; Zhang et al., 2011).

Evaluating the topographic-atmospheric feedbacks introduced here is well beyond the scope of this study. Our analysis has provided a key boundary condition for climate modelers to incorporate in their simulations of Paleogene Earth, i.e., the paleolatitude of the southern margin of Asia and the proto-Tibetan Plateau.

CONCLUSIONS

Our compilation of diverse paleomagnetic records from the Lhasa terrane has been evaluated in a statistically consistent framework to calculate a paleolatitude record for the southern margin of the terrane since ca. 110 Ma. Our analysis suggests that the southern margin of Asia has remained relatively stable with respect to the Earth's spin axis throughout the Cretaceous and early Paleogene, and has moved northward by ~8° of latitude mostly since 50 Ma. The paleomagnetically estimated collision age between the Tibetan Himalaya and the southern margin of Asia is ca. 49.5 ± 4.5 Ma, if not a few million years earlier. Adding reasonable estimates for the width of the original Asian forearc, and subducted northern Tibetan Himalayan passive margin rocks may revise this collision age to ca. 52 ± 5 Ma. These age estimates are in excellent agreement with stratigraphic records of this collisional event suggesting a ca. 52 Ma collision. This collision occurred at lat $\sim 21^\circ \pm 4^\circ N$, or perhaps $\sim 2^\circ$ lower if a reasonably sized forearc is added. A Tibetan Himalayan-Lhasa collision ca. 52-50 Ma at ~21°N indicates that at most only $\sim 1100 \pm 560$ km of India-Asia convergence was partitioned into Asian lithosphere. The low end of these paleomagnetic estimates is consistent with the ~740 km of upper crustal shortening within Asia calculated from structural geologic studies (van Hinsbergen et al., 2011a). A direct implication of these findings is that 1700 ± 560 km or more convergence was partitioned into the lower plate (i.e., units of Indian affinity). Thus, Greater India at the time of the Tibetan Himalaya-Lhasa collision was much larger than constrained by pre-Early Cretaceous paleomagnetic data and tectonic reconstructions. As considered elsewhere (van Hinsbergen et al., 2012b), these observations require significant extension of Greater India lithosphere after break-up from Gondwana but prior to collision with the southern margin of Asia. Cretaceous extension within Greater India resulted in the oceanic Greater India Basin, which may have maintained oceanic circulation between the Equatorial Pacific and eastern Mediterranean Tethys throughout most of the Paleogene. The significance of this marine gateway to early Cenozoic climate dynamics and surface processes within the Tethyan realm is beyond the scope of this paper, but we suspect it may be large. Our analysis suggests that the southern margin of the Tibetan Plateau has been positioned at high tropical to low subtropical latitudes since at least the end of Early Cretaceous. Future climate model simulations should incorporate this paleogeography to more accurately explore how atmospheric processes interact with or are modified by the high uniform topography of the Tibetan Plateau. Teleconnections between atmospheric jets, regional precipitation patterns, and the long-term carbon cycle may be particularly sensitive to the paleolatitude of orogenic plateaus.

ACKNOWLEDGMENTS

Lippert was supported by National Science Foundation Continental Dynamics grant EAR-1008527; van Hinsbergen acknowledges financial support from Statoil (SPlates Model Project), the Center for Advanced Study of the Norwegian Academy of Science and Letters, and European Research Council Starting Grant SINK (Subduction Initiation Reconstructed from Neotethyan Kinematics; project no. 306810); and Dupont-Nivet acknowledges support from the Netherlands Organization for Scientific Research, the Centre National de la Recherche Scientifique, and the National Science Foundation of China. We thank Paul Kapp and Joellen Russell for discussions on parts of this manuscript, and John Geissman and Erwin Appel for thorough reviews.

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