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# Triassic (Anisian and Rhaetian) palaeomagnetic poles from the Germanic Basin (Winterswijk, the Netherlands)

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## Abstract

In this paper, we provide two new Triassic palaeomagnetic poles from Winterswijk, the Netherlands, in the stable interior of the Eurasian plate. They were respectively collected from the Anisian (~247–242 Ma) red marly limestones of the sedimentary transition of the Buntsandstein Formation to the dark grey limestones of the basal Muschelkalk Formation, and from the Rhaetian (~208–201 Ma) shallow marine claystones that unconformably overlie the Muschelkalk Formation. The magnetization is carried by hematite or magnetite in the Anisian limestones, and iron sulfides and magnetite in the Rhaetian sedimentary rocks, revealing for both a large normal polarity overprint with a recent (geocentric axial dipole field) direction at the present latitude of the locality. Alternating field and thermal demagnetization occasionally reveal a stable magnetization decaying towards the origin, interpreted as the Characteristic Remanent Magnetization. Where we find a pervasive (normal polarity) overprint, we can often still determine well-defined great-circle solutions. Our interpreted palaeomagnetic poles include the great-circle solutions. The Anisian magnetic pole has declination  $D \pm \Delta D_x = 210.8 \pm 3.0^\circ$ , inclination  $I \pm \Delta I_x = -26.7 \pm 4.9^\circ$ , with a latitude, longitude of  $45.0^\circ$ ,  $142.0^\circ$  respectively,  $K = 43.9$ ,  $A_{95} = 2.9^\circ$ ,  $N = 56$ . The Rhaetian magnetic pole has declination  $D \pm \Delta D_x = 32.0 \pm 8.7^\circ$ , inclination  $I \pm \Delta I_x = 50.9 \pm 8.1^\circ$ , with a latitude, longitude of  $60.6^\circ$ ,  $123.9^\circ$  respectively,  $K = 19.3$ ,  $A_{95} = 7.4^\circ$ ,  $N = 21$ . The poles plot close to the predicted location of global apparent polar wander paths (GAPWaps) in Eurasian coordinates and are feasible for future apparent polar wander path construction. They confirm that the intracontinental, shallow-marine Germanic Basin, in which the Muschelkalk Formation was deposited, existed at a palaeolatitude of  $14.1^\circ$  [11.3, 17.1] N, in a palaeo-environment reminding of the Persian Gulf today. In Rhaetian times, palaeolatitudes of  $31.6^\circ$  [24.8, 39.8] N were reached, on its way to the modern latitude of  $52^\circ$ N.

**Keywords:** Palaeomagnetic, Palaeolatitude, Anisian, Rhaetian, Muschelkalk Formation, Germanic Basin

## 1 Introduction

Palaeogeographic reconstructions provide essential context and boundary conditions for the study of deep-time geological, climatological, and biological processes (Blakey 2008; Scotese and Schettino 2017; Torsvik and Cocks 2017). For times of the supercontinent Pangea, in Triassic time (~251–201 Ma), such reconstructions are primarily based on plate tectonic reconstructions using marine magnetic anomalies, which show how continents

and oceans have moved relative to each other (e.g., Seton et al. 2012). Subsequently, plate margins are corrected for intense plate boundary zone deformation using structural geological constraints (e.g., Hall 2002; Boschman et al. 2014; van Hinsbergen et al. 2014, 2019). To these reconstructions, geographical features such as coastlines, mountain belts, lakes and seas, and icecaps are then added using geological data (Vérard et al. 2015; Baatsen et al. 2016; Scotese and Schettino 2017; Torsvik and Cocks 2017). Finally, to use the reconstruction as basis for wider geoscientific study, the reconstruction needs to be placed in a reference frame that determines the position of the plates, continents, and oceans relative

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to the Earth's spin axis (for palaeo-climate and environment studies), or relative to modern mantle structure (for studies on geodynamic processes) (e.g., Cox and Hart 1986; Besse and Courtillot 2002; Torsvik et al. 2008; Kent and Irving 2010; van Hinsbergen et al. 2015).

On geological timescales, the Earth's magnetic field may be described as dipolar, aligning along the spin axis (e.g., Tauxe 2010). The positions of plates relative to the poles are then reconstructed by measuring the palaeomagnetic field recorded by and preserved in rocks. From the direction of the magnetic field preserved in a rock, a location may be calculated of the associated palaeomagnetic pole, and when that pole does not coincide with the modern geographic north pole, net plate motion relative to such poles may be calculated (e.g., Tauxe 2010). With relative motions of major plates relative to each other constrained, all poles from stable plate interiors may be rotated into a common reference location, and a global apparent polar wander path (GAPWaP) may be calculated (e.g., Besse and Courtillot 2002; Kent and Irving 2010; Torsvik et al. 2012). A GAPWaP provides the position of a palaeogeographic reconstruction of plates and continents relative to the Earth's spin axis, and forms the reference for studies applying palaeomagnetism to reconstruct intensely deformed plate boundary zones, or to reconstruct plates from intensely deformed relics accreted at plate boundaries (e.g., Nokleberg et al. 2000). Whereas applications of the available GAPWaP are numerous, and palaeomagnetic datasets aiming to constrain regional deformation relative to major plates rapidly increase in size (e.g., Cogné et al. 2013; Li et al. 2017), new data from tectonically quiet plate interiors are sparse, and the resolution of the reference GAPWaP may become a limiting factor (Dupont-Nivet et al. 2010). It is therefore important to collect new poles from stable continental interiors.

In this paper, we provide such a new pole, from the Netherlands. During the Anisian, the area of the Netherlands was located in the Germanic Basin in the heart of Pangea (Ziegler 1992). During the Early Triassic, a regional transgression drowned a widespread fluvio-lacustrine basin (Buntsandstein), to give way to the deposition of a shallow-marine carbonate formation known as the Muschelkalk Formation (e.g., Ziegler 1992). Deformation in the region was restricted to some normal faulting, and palaeomagnetic poles from Northwest Europe are widely considered representative for stable Eurasia (e.g., Torsvik et al. 2012). In the Netherlands, the Buntsandstein-Muschelkalk sedimentary environment transition is exposed in the quarries of Winterswijk (Fig. 1), and is dated as the Anisian (Herngreen et al. 2005) (Middle Triassic, 247.2–242 Ma; Gradstein et al. 2012). Here we report palaeomagnetic and rock magnetic analyses of rocks from red and white marly

limestones and carbonates from the Anisian uppermost Buntsandstein Formation and lowermost Muschelkalk, and from overlying Rhaetian dark grey claystones, to constrain new palaeomagnetic poles, and provide a first-order reconstruction displaying the position of the Winterswijk deposits in their palaeogeographic context.

## 2 Geological setting and sampling

The 'Winterswijkse Steen- en Kalkgroeve' (51.97°N, 6.78°E) comprises a series of quarries near Winterswijk, in the east of the Netherlands (Fig. 1), where the lower Muschelkalk Formation is mined for building material. The base of the quarry exposes a ~3°–9° dipping monoclinical succession, consisting of ~40 m thick section comprising the top few meters of Buntsandstein red marly limestone, followed by a transition of paralic Muschelkalk blue-grey marly limestones to the shallow marine (Herngreen et al. 2005) (Fig. 1). This well-dated section, palynologically constrained at the Anisian (247.2–242 Ma; Herngreen et al. 2005), has a rich fossil fauna, which has been subject to numerous palaeontological studies. These recovered ichnofossils and tetrapod footprints (Oosterink 1976; Demathieu and Oosterink 1988; Diedrich 2000; Schulp et al. 2017), reptiles (nothosaurs and placodonts) (Albat 1999; Bickelmann and Sander 2008; Klein 2009, 2012; Klein and Albers 2009; Albers 2011; Klein and Scheyer 2013; Klein and Sichelschmidt 2014; Sander et al. 2014; Klein et al. 2015, 2016; Voeten et al. 2015, 2018; Maxwell et al. 2016; During et al. 2017; Klein and Griebeler 2018; Heijne et al. 2019), but also lobsters (Klomp maker and Fraaije 2011), jellyfish (Klomp maker and Fraaije 2011), and insects (van Eldijk et al. 2017). The Muschelkalk Formation contains widespread secondary mineralizations in cracks and pores, containing like pyrite and celestine (Kloprogge et al. 2001).

The Muschelkalk Formation is unconformably covered by the uppermost Triassic (Rhaetian, 208.5–201.3 Ma; Gradstein et al. 2012), pyrite-bearing, poorly consolidated claystone (Herngreen et al. 2005). Besides, subsidence pipes (sinkholes) were found within the Muschelkalk, interpreted to a result from the collapse of subsurface cavities that are thought to have formed due to the dissolution of Zechstein salts of the underlying Permian sequence (e.g., Klomp maker et al. 2010). These are filled with the Rhaetian-Hettangian (the uppermost Triassic to the lowermost Jurassic) ammonite-bearing claystone (Herngreen et al. 2005; Oosterink et al. 2006; Klomp maker and Van den Berkmortel 2007; Klomp maker et al. 2010).

We collected 222 palaeomagnetic core samples from nine stratigraphic horizons (WW 1–9) across the 40 m of the Muschelkalk Formation: 80 from four horizons of red marly limestone (WW 1, 5–7) (Fig. 1d), and 142 from grey limestone (WW 2–4, 8–9) (Fig. 1b, c). At each stratigraphic horizon, 20–40 samples were collected across an approximately 1.5 m thick interval, which with



**Fig. 1** **a** - Map of Europe with location of the Winterswijk sampling site in the east of the Netherlands; **b-c** - Field photographs of the sampled Muschelkalk Formation limestones; **d** - Field photograph of the sampled red marly limestones interlayered with the Muschelkalk Formation limestones; **e** - Field photograph of the Rhaetian Formation claystones, which were so soft that they were sampled with plastic tubes, here sticking out of the sediments

typical sedimentation rates of several to tens of cm/kyr is sufficient to assume the palaeomagnetic direction scatter obtained from these horizons or from all horizons combined and represent the secular variation of the geomagnetic field. Sampling occurred with a water-cooled, generator-powered electric drill, and samples were oriented with a magnetic compass. Orientations needed no correction for the local declination ( $0^\circ$  at the time of sampling). In addition, we collected 73 samples from a  $\sim 1$  m thick section (WW 10) of Rhaetian claystones unconformably overlying the Muschelkalk Formation on the north side of the quarry. Because these clays were poorly consolidated and disintegrated during sampling with a water-cooled drill, we collected samples using 25 mm diameter plastic tubes that were hammered into the clays (Fig. 1e), a common sampling technique for poorly consolidated sediments (e.g., Sier et al. 2011). The benefit of the tubes was that they have the same size as regular palaeomagnetic cores and could in the field be oriented with a magnetic compass and extracted from the exposure.

### 3 Methods

Cores and sample tubes were cut into standard specimens of 25 mm diameter and 22 mm length. At least one specimen per core was demagnetized. We carried out the laboratory analyses at Paleomagnetic Laboratory Fort Hoofddijk, Utrecht University, the Netherlands.

The natural remanent magnetization (NRM) of at least one specimen per core or tube sample was progressively demagnetized according to standard stepwise palaeomagnetic practice (Zijderveld 1967). Thermal demagnetization was done with steps of 20–50 °C up to 680 °C using an ASC thermal demagnetizer (model TD48, residual field < 10 nT) and remaining NRMs were measured with a '2G' enterprises DC SQUID (Direct Current Superconducting Quantum Interference Device) magnetometer with a noise level of  $3 \times 10^{-12}$  Am<sup>2</sup>. Alternating field (AF) demagnetization was done with steps of 5–10 mT up to 100 mT using the robotized demagnetization setup at Fort Hoofddijk (Mullender et al. 2016). The tube samples from the Rhaetian claystones were only AF demagnetized as they were contained in a plastic tube. A total of 192 specimens from cores of the Anisian red marly limestones and grey limestones of the lower Muschelkalk were AF demagnetized and an additional 48 specimens were demagnetized thermally (see Additional files 1 and 2).

Demagnetization diagrams were plotted on orthogonal vector diagrams (Zijderveld 1967), and interpreted via principal component analysis (Kirschvink 1980) as criterion for determining the linearity or planarity of demagnetization data (when a single directional component is being demagnetized forming a linear segment, or

when two coercivity or blocking temperature spectra overlap forming great circles demagnetization paths) in terms of Maximum Angular Deviation (MAD) values. Mean directions for the Anisian and Rhaetian samples were calculated combining lines and converging planes based on all available data for each stratigraphic interval (McFadden and McElhinny 1988). In fitting the great circles, we determine the modified  $\alpha_{95}$  ( $\alpha_{m95}$ ) for setpoints and great circle solutions, as given by McFadden and McElhinny (1988). Paleomagnetism.org also provides the full  $\alpha_{95}$  as derived from their equation (25). Directional means, as well as means of the virtual geomagnetic poles (VGPs), were calculated according to Fisher (1953), including their corresponding dispersion and cones of confidence ( $k$ ,  $\alpha_{95}$  and  $K$ ,  $A_{95}$ , respectively). Errors in declination ( $\Delta D_x$ ) and inclination ( $\Delta I_x$ ) were determined from  $A_{95}$  according to Butler (1992). We used the criteria of Deenen et al. (2011, 2014) to determine whether secular variation has been sufficiently sampled, based on the N-dependent envelope of  $A_{95}$  ( $A_{95\min}$ – $A_{95\max}$ ). A fixed cut-off ( $45^\circ$ ) was applied to remove outliers. For interpretation and statistical treatment, we used the online palaeomagnetic analysis platform [Paleomagnetism.org](http://Paleomagnetism.org) (Koymans et al. 2016). Files compatible with this platform containing all data and interpretations are provided in the Additional files 1 and 2.

To determine the nature of the magnetic carriers, we performed rock magnetic analyses including thermomagnetic runs (5 samples) with a modified horizontal translation-type Curie balance with a cycling applied magnetic field of 150–300 mT (Mullender et al. 1993). To discriminate between thermochemical alteration during the experiment and truly magnetic behavior, we cycled to increasingly elevated temperatures with cooling segments in between, 250 °C, 150 °C, 350 °C, 250 °C, 450 °C, 350 °C, 520 °C, 420 °C, 620 °C, 500 °C, 700 °C; the entire sequence was room temperature. In addition, we measured detailed acquisition curves of the isothermal remanent magnetization (IRM) up to 700 mT (61 IRM field levels; 52 specimens) using the robotized 2G Enterprises SQUID magnetometer with in-line alternating field (AF) demagnetization and IRM acquisition facilities (Mullender et al. 2016). The latter is a laboratory-induced remanence that is used to characterize the magnetic minerals carrying the NRM along with their grain size that further serves to make inferences concerning the palaeomagnetic stability of the NRM. IRM acquisition curves were obtained from the so-called static 3-axes AF demagnetized starting state, with the final AF demagnetization axis parallel to the subsequent IRM (pulse) fields. In this way, there is a minimal deviation in the shape of the measured IRM acquisition curve from a cumulative log-Gaussian distribution (cf. Heslop et al. 2004). Thus, before the IRM acquisition, we demagnetized the samples along three orthogonal axes at 300 mT AF using a laboratory-built

demagnetization coil. IRM coercivity fraction fitting was done with the interactive software package of Kruiver et al. (2001) that assumes cumulative log-Gaussian IRM acquisition curves are linearly additive. Each coercivity fraction is determined by three parameters: SIRM (saturation IRM; the amount of the specific magnetic mineral),  $B_{1/2}$  (the IRM acquisition field at which 50% of the IRM is reached) and DP (dispersion parameter, describing the width of an IRM coercivity distribution). An IRM acquisition curve is the sum of several of such coercivity components. In reality, IRM acquisition curves are often not completely log-Gaussian, which is a consequence of thermal activation, in particular, fine-grained particles close to the superparamagnetic threshold (~ 30 nm at room temperature), and show time-dependent behavior. This results in “skewed-to-the-left” coercivity distributions (Egli 2004; Heslop et al. 2004) which require an additional coercivity component for a proper fitting in the Kruiver et al. (2001) package. In the interpretation, the amount of skew-component is added to that of the soft component, representing magnetite (*sensu lato*, thus including partially oxidized magnetite, i.e., the range of compositions between  $Fe_3O_4$  and  $\gamma-Fe_2O_3$ ). The hard IRM component then represents hematite ( $\alpha-Fe_2O_3$ ).

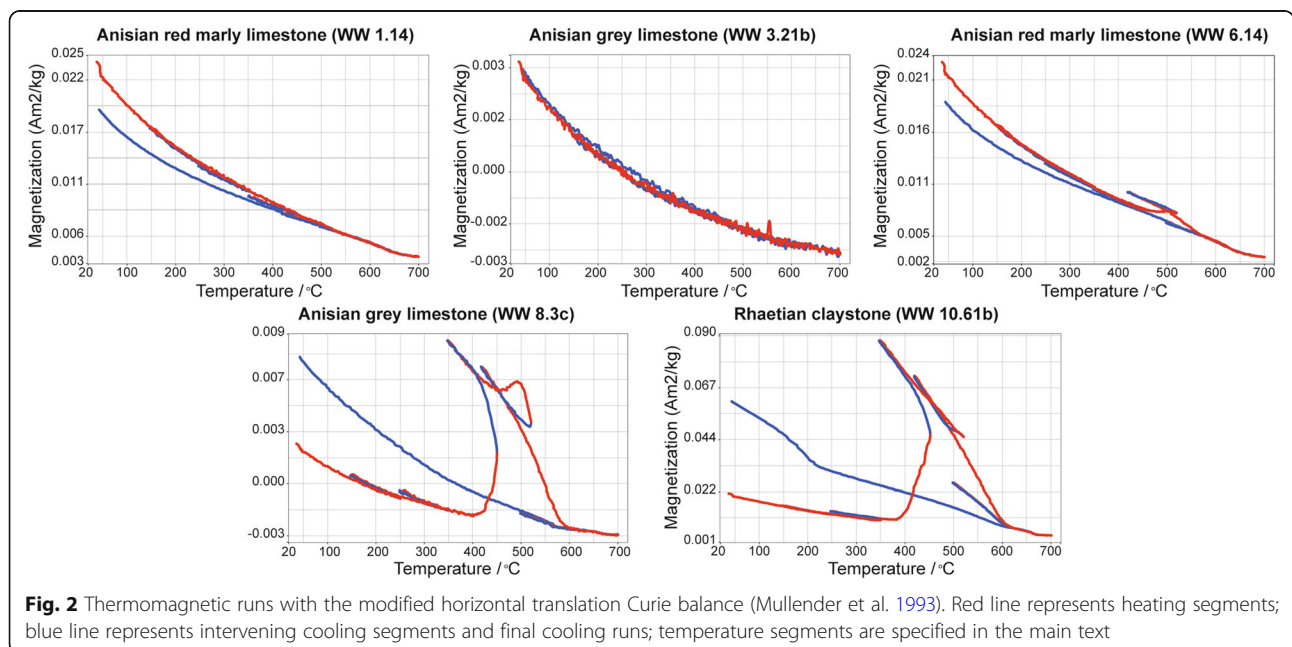
## 4 Results

### 4.1 Rock magnetism

#### 4.1.1 Thermomagnetic analysis

Thermomagnetic runs indicate that samples are only weakly magnetic as expected for these types of sediments (Fig. 2). WW 1.14 and WW 6.14 are from the Anisian red marly limestones. They are essentially paramagnetic; during cycling

above 500 °C some magnetization is lost but a smaller magnetization emerges during the final cooling runs. Both samples show a minutely expressed Néel point of hematite. WW 6.14 shows some reaction to magnetite starting at ~ 430 °C; presumably, during the thermomagnetic run, pyrite is being oxidized via magnetite to hematite. WW 3.21b and WW 8.3c are from the Anisian grey limestones. They are weakly magnetic only with WW 3.21b showing no measurable ferromagnetism. In WW 8.3c the oxidation reaction of pyrite is markedly present. Apparently, both red and grey limestones have variable amounts of pyrite. The sample WW 10.61b from the Rhaetian claystone, is fully reversible to 250 °C. The small ‘hump’ at 300–350 °C may indicate ferrihydrite conversion to magnetite due to the reducing capacity of organic matter, which then converts to hematite at higher temperatures. The increase starting at ~ 400 °C is interpreted to be caused by pyrite oxidation, again first to magnetite and finally to hematite (e.g., Passier et al. 2001). Hematite’s presence is clear from its Néel point at ~ 675 °C. Magnetite is completely oxidized to hematite at elevated temperature since a Curie point does not reappear during the final cooling run (from 700 °C to room temperature). It should be noted that the negative magnetic moment at high temperature in some samples due to diamagnetism of the quartz glass sample holder. The magnetite formation during the thermomagnetic experiment has a direct bearing on thermal demagnetization of the NRM: it does not bode well for samples that contain traces of pyrite. The impossibility to thermally demagnetize the samples because they were contained in plastic tubes, therefore, was of little influence: the spurious magnetic moment of the formed magnetite would have completely obscured the NRM

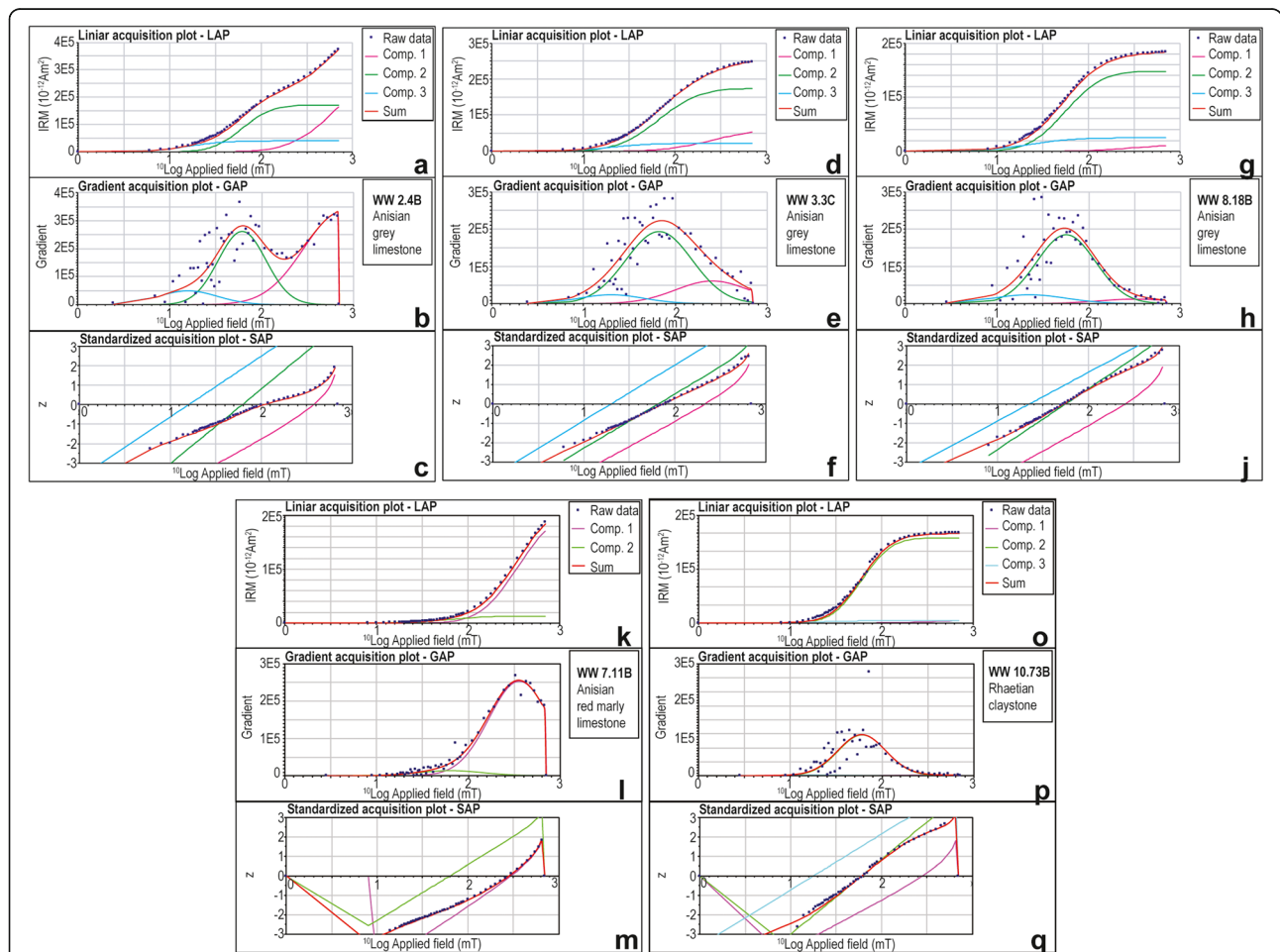


during thermal demagnetization steps above ~ 450 °C. This is why the Rhaetian samples were treated with AF demagnetization solely.

**4.1.2 IRM acquisition curves**

The IRM acquisition curves are fitted with three components of which two are deemed meaningful and interpreted (Additional file 1: Supplementary Information 1a). Component 1 with  $B_{1/2}$  ranging from ~ 150 mT to over 1 T, typically between 250 mT and 350 mT, is interpreted as hematite. Samples with a high contribution of this component are often reddish colored. Component 2 with  $B_{1/2}$  ranging from ~ 50 mT to ~ 70 mT, often in between ~ 55 mT and ~ 65 mT, is interpreted as magnetite oxidized to a variable

degree. The third, lowest field, component (component 3; Additional files 1 and 2) is inferred to be the consequence of thermal activation; its amount is added to the component 2. It is inherent to the symmetric distributions in the log-Gaussian field space enforced by the fitting package. Below we illustrate the IRM acquisition behavior (Fig. 3). Sample WW 2.4B (Fig. 3) from the Anisian grey limestone has an appreciable hematite contribution (62.5% of the total (calculated) IRM). It is magnetically notably hard with a  $B_{1/2}$  of ~ 760 mT; its dispersion parameter of 0.42 (all DP are in log units) is fairly high but not uncommon for an antiferromagnetic material like hematite. The magnetite amount, added components 2 and 3, is 37.5%; it has a  $B_{1/2}$  of ~ 62 mT and a dispersion



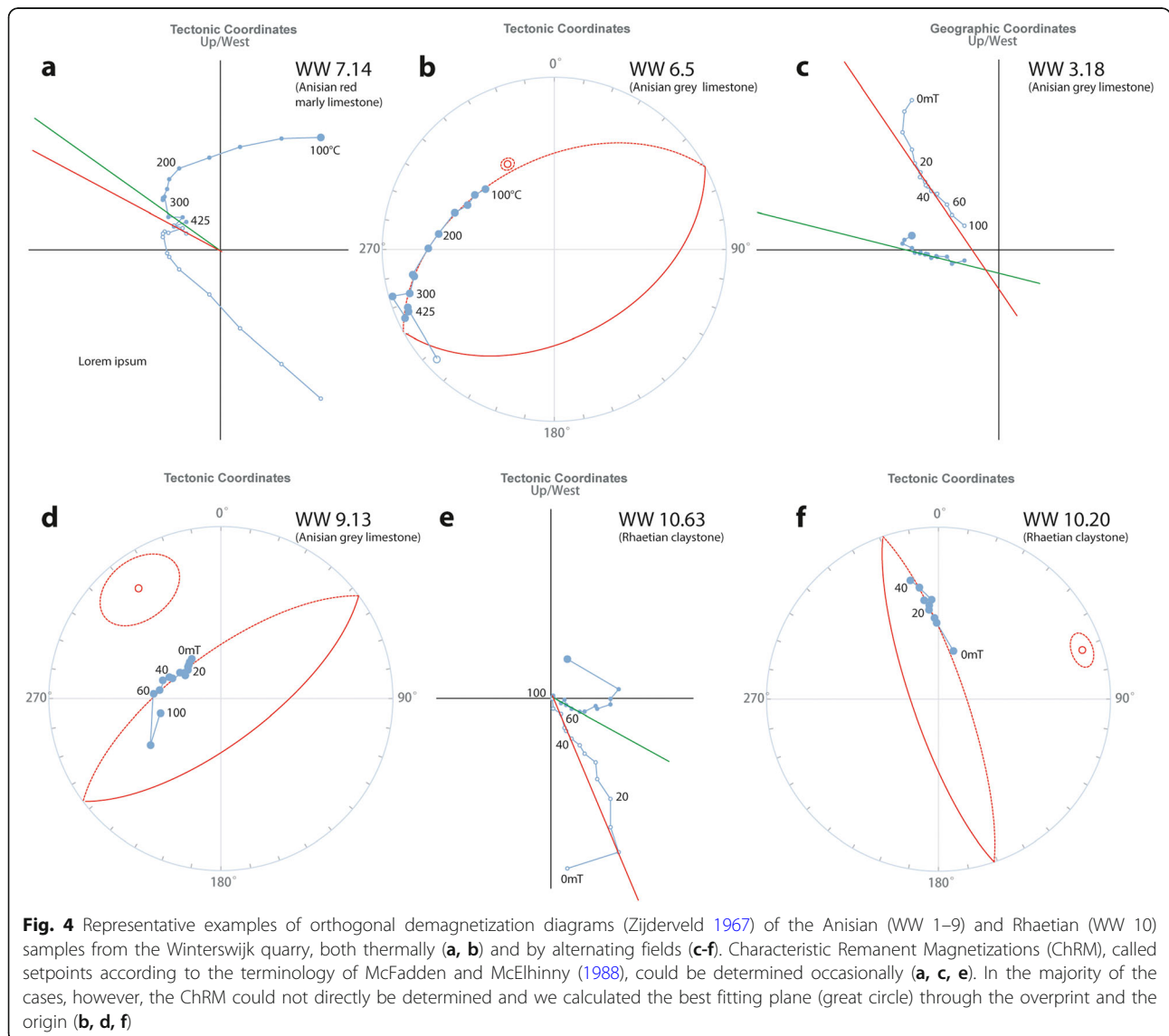
**Fig. 3** Examples of coercivity component fits to IRM acquisition curves. **a-c** - Sample WW 2.4B from the Anisian grey limestone; **d-f** - Sample WW 3.3C from the Anisian grey limestone; **g, h, j** - Sample WW 8.18B from the Anisian grey limestone; **k-m** - Sample WW 7.11B from the Anisian red marly limestone; **n-q** - Sample WW 10.73B from the Rhaetian claystone. Component (Comp.) 1 (pink) is interpreted as hematite, Comp. 2 (green) is variably oxidized magnetite, and Comp. 3 (light blue) is the 'skew' component, see main text for detailed explanation. The sum of the components is marked in red; data points are the grey points. In the linear acquisition plots (LAP), i.e., **a, d, g, k, o**, the magnetic moment is plotted versus the logarithmic field value; in the gradient acquisition plots (GAP), i.e., **b, e, h, l, p**, the gradient of the LAP is plotted; in the standardized acquisition plots (SAP), i.e., **c, f, j, m, q**, the standardized score ( $z$ ) is plotted versus the logarithmic field value. In this way of plotting deviations from cumulative log-Gaussian distributions show up as deviations from a straight line, indicating the need of additional components for a good fit

parameter of 0.26 which is in the often-observed range for relatively non-oxidized detrital magnetite (Kruiver et al. 2003). Magnetotactic magnetite would have a higher  $B_{1/2}$  and a notably smaller DP, below 0.18 or even lower (e.g., Kruiver and Passier 2001; Egli 2004). The sample appears to have a reasonably low hematite amount; other samples show hematite amounts up to two orders of magnitude higher (Additional file 1: Supplementary Information 1b).

Sample WW 3.3C (Fig. 3) from the Anisian grey limestone contains 23.7% hematite with a lower  $B_{1/2}$  of ~ 250 mT and a dispersion parameter of 0.4. When the relative amount of hematite decreases, it tends to be magnetically softer, hinting at a finer grain size. The magnetite (76.3%) is partially oxidized, in principle raising the  $B_{1/2}$  value (here  $B_{1/2}$  is ~ 66 mT) and the dispersion parameter value

(0.36 in this sample). The dispersion parameter value is notably high for a ferrimagnetic material, indicating a broad switching field distribution which points at partially oxidized magnetite (Kruiver et al. 2003). Sample WW 8.18B from the Anisian grey limestone (Fig. 3) contains only 7% hematite with similar properties as in sample WW 3.3C. The dominant magnetite (93%) is fairly soft ( $B_{1/2}$  ~ 56 mT) with a moderate dispersion parameter of 0.32. The thermally activated portion (component 3) seems to be higher in such samples.

Sample WW 7.11B also from the Anisian red marly limestone (Fig. 3) is evidently dominated by hematite with a hematite component representing almost 95% of the total IRM. It is relatively soft with the  $B_{1/2}$  value being 355 mT and the dispersion parameter of 0.33. The minute amount of magnetite is probably oxidized with



**Fig. 4** Representative examples of orthogonal demagnetization diagrams (Zijderveld 1967) of the Anisian (WW 1–9) and Rhaetian (WW 10) samples from the Winterswijk quarry, both thermally (a, b) and by alternating fields (c–f). Characteristic Remanent Magnetizations (ChRM), called setpoints according to the terminology of McFadden and McElhinny (1988), could be determined occasionally (a, c, e). In the majority of the cases, however, the ChRM could not directly be determined and we calculated the best fitting plane (great circle) through the overprint and the origin (b, d, f)

its comparatively high  $B_{1/2}$  of 63 mT and high dispersion parameter of 0.35.

Sample WW 10.73B from the Rhaetian claystone (Fig. 3) is magnetically dominated by magnetite (components 2 and 3 add up to 96.6%) with just about 3% as the hematite contribution.  $B_{1/2}$  of magnetite is ~ 60 mT, together with the relatively low dispersion parameter (0.26), indicating non-oxidized magnetite.  $B_{1/2}$  of hematite is ~ 400 mT, with the dispersion parameter of 0.42, indicating a broad grain size distribution.

Two groups of samples are distinguished: one with a low hematite content of < 40% (from a magnetic viewpoint) and another with a high hematite content of > 40% (Additional file 1: Supplementary Information 1a). The first group ( $N = 44$ ) has an average of ~ 14% hematite, and the second group ( $N = 8$ ) has an average of ~ 67% hematite. When the hematite amount is low, hematite appears to be magnetically softer: its average  $B_{1/2}$  value is ~ 270 mT. In hematite-dominated samples,  $B_{1/2}$  values are distinctly higher (average  $B_{1/2}$  is ~ 610 mT). The magnetite in the low-hematite group is slightly softer as well with a  $B_{1/2}$  of ~ 58 mT compared to ~ 61 mT for the high-hematite group. The standard deviations of both groups, however, largely overlap. Therefore, we recognize one type of magnetite throughout the studied section.

## 4.2 Demagnetization

### 4.2.1 Anisian

Both thermal and AF demagnetization of the Anisian samples (WW 1–9, see Additional files 1 and 2 for data files) show a large overprint of normal polarity with an approximately recent field (normal) direction which is likely due to the weathering of the rocks. It was not easily removed, and in only 10 samples we could determine

a ChRM direction as a component which was of reversed polarity and decayed to the origin, so called setpoints (e.g., WW 7.14 from red marly limestone, Fig. 4a). The normal component dominates and the fully reversed direction that we find in the ChRM is not or not fully reached in all other samples. In that case, we determine a great circle (plane), fitting a selected arc of demagnetization data, assuming (1) there is no more than 2 components being simultaneously demagnetized in the arc, (2) the original component is contained in the plane but outside the selected demagnetization arc (e.g., WW 6.5 from grey limestone, Fig. 4b). AF demagnetization does not differ significantly from the thermal treatment, and we find both setpoints and great circles that are compatible with the thermal results (WW 3.18 and WW 9.13, both from grey limestones, Fig. 4c, d). We can now determine the mean directions (Table 1) either based on setpoints only ( $sp$ ;  $N = 13$ ) and/or based on combining setpoints and great circles ( $sp + gc$ ;  $N = 56$ ) using the iteration method to find the best-fit great circle solutions (McFadden and McElhinny 1988); the results of the two distributions are shown in Fig. 5. The means differ in declination (from  $D_{sp} = 40.9 \pm 8.0^\circ$  to  $D_{sp + gc} = 30.8 \pm 3.0^\circ$ ), and, not surprisingly, the error in declination ( $\Delta D_x$ ) becomes smaller with a larger number of samples in the two distributions. The difference in inclination ( $0.3^\circ$ ), however, is negligible (Table 1). Therefore, the test of the two means of sharing the same distribution using the bootstrap coordinate test (Tauxe 2010) is negative (Fig. 6).

### 4.2.2 Rhaetian

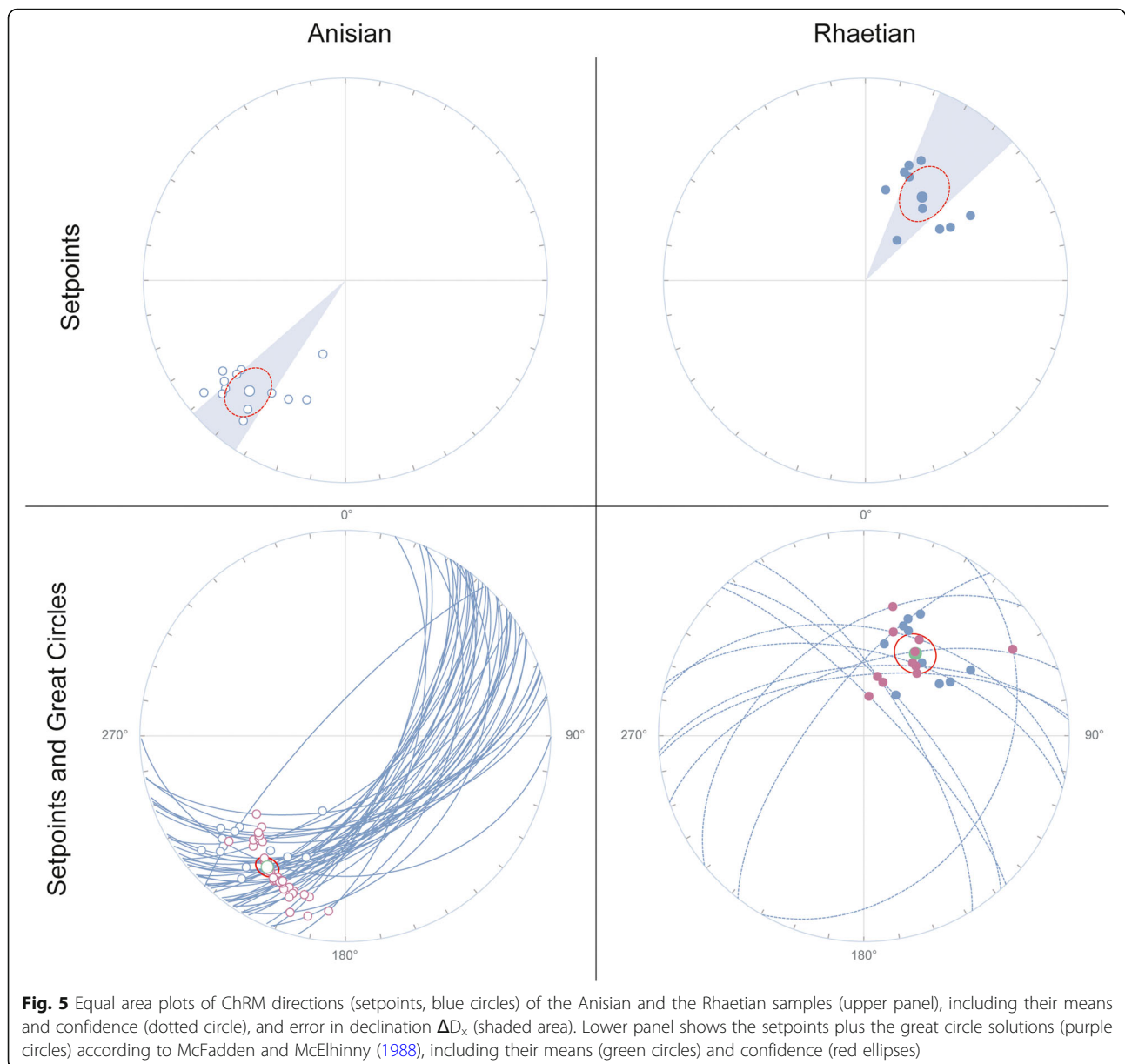
AF demagnetization of the Rhaetian samples (WW 10) shows a large overprint of normal polarity with an

**Table 1** Palaeomagnetic data obtained from the Winterswijk quarry. Sites (site names) indicated with '(sp)' only include ChRM directions obtained by fitting linear segments of demagnetization paths (setpoints) interpreted using principal component analysis (Kirschvink 1980). Sites (site names) indicated with '(sp + gc)' also include great circle solutions referring to McFadden and McElhinny (1988)

Site Names	$N_s$	$N_{45}$	Dec	$\Delta D_x$	Inc	$\Delta I_x$	K	$A_{95min}$	$A_{95}$	$A_{95max}$	pLat	pLong	$\lambda_{lo}$	$\lambda$	$\lambda_{hi}$
Geographic Coordinates															
Rhaetian (sp)	10	10	43.3	± 14.6	54.0	± 12.1	17.3	4.8	< 11.9	< 19.2	-	-	-	-	-
Rhaetian (sp + gc)	21	21	41.7	± 10.1	56.5	± 7.7	16.5	3.6	< 8.1	< 12.0	-	-	-	-	-
Anisian (sp)	13	13	224.5	± 8.4	-31.7	± 12.7	27.6	4.3	< 8.0	< 16.3	-	-	-	-	-
Anisian (sp + gc)	56	56	214.8	± 3.2	-31.3	± 4.9	39.4	2.4	< 3.1	< 6.5	-	-	-	-	-
Tectonic Coordinates															
Rhaetian (sp)	10	10	34.2	± 12.6	48.7	± 12.6	20.5	4.8	< 10.9	< 19.2	56.9	123.4	20.0	< 29.7	< 42
Rhaetian (sp + gc)	21	21	32.0	± 8.7	50.9	± 8.1	19.3	3.6	< 7.4	< 12.0	60.6	123.9	24.8	< 31.6	< 39.8
Anisian (sp)	13	13	220.9	± 8.0	-28.4	± 12.8	30.3	4.3	< 7.7	< 16.3	41.9	130.2	8.0	< 15.1	< 24
Anisian (sp + gc)	56	56	210.8	± 3.0	-26.7	± 4.9	43.9	2.4	< 2.9	< 6.5	45.0	142.0	11.3	< 14.1	< 17.1

Key to abbreviations:  $N_s$  = Number of interpreted ChRM directions;  $N_{45}$  = Number of directions after applying a 45° cutoff; Dec = Declination;  $\Delta D_x$  ( $\Delta I_x$ ) = 95% confidence interval on the declination (inclination) following Butler (1992); Inc = Inclination; K = Fisher (1953) precision parameter calculated on virtual geomagnetic poles (VGPs);  $A_{95}$  = 95% cone of confidence around the VGPs;  $A_{95(min, max)}$  = Reliability envelope of Deenen et al. (2011); pLat, pLong = Pole latitude and longitude;  $\lambda$  = Palaeolatitude;  $\lambda_{(lo, hi)}$  = Lower and upper confidence limits for paleolatitude





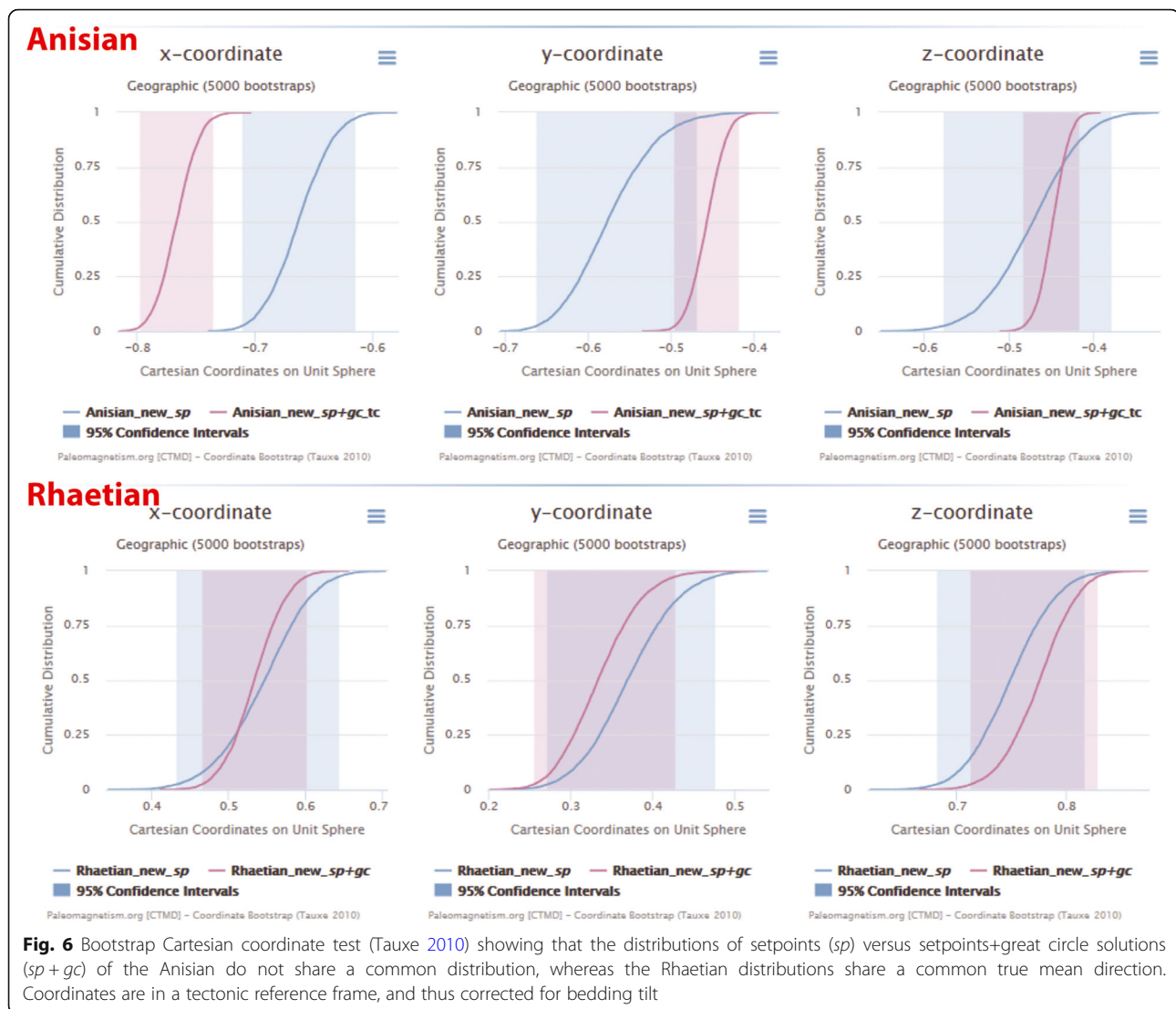
approximately recent field (normal) direction. The overprint is also not easily removed and in only 10 samples we determine setpoints, of normal polarity, that decayed towards the origin (e.g., WW 10.63, Fig. 4e). In most cases, we could only determine a great circle (e.g., WW 10.20, Fig. 4f). We could only use AF demagnetization, so we cannot compare results with those from thermal treatment. We determine the mean directions (Table 1) based both on setpoints only ( $sp$ ;  $N = 10$ ) and on combining setpoints and great circle solutions ( $sp + gc$ ;  $N = 21$ ). The means of the two distributions do not differ significantly within error, neither in declination (from  $D_{sp} = 34.2 \pm 12.6^\circ$  to  $D_{sp+gc} = 32.0 \pm 8.7^\circ$ ) nor in inclination (differing

by  $0.1^\circ$ ; Table 1). In the case of the Rhaetian samples, the test of the two means of sharing the same distribution is positive (Tauxe 2010; Fig. 6), so the two means can be considered indistinguishable.

## 5 Discussion

### 5.1 Palaeomagnetic poles for the Anisian and the Rhaetian

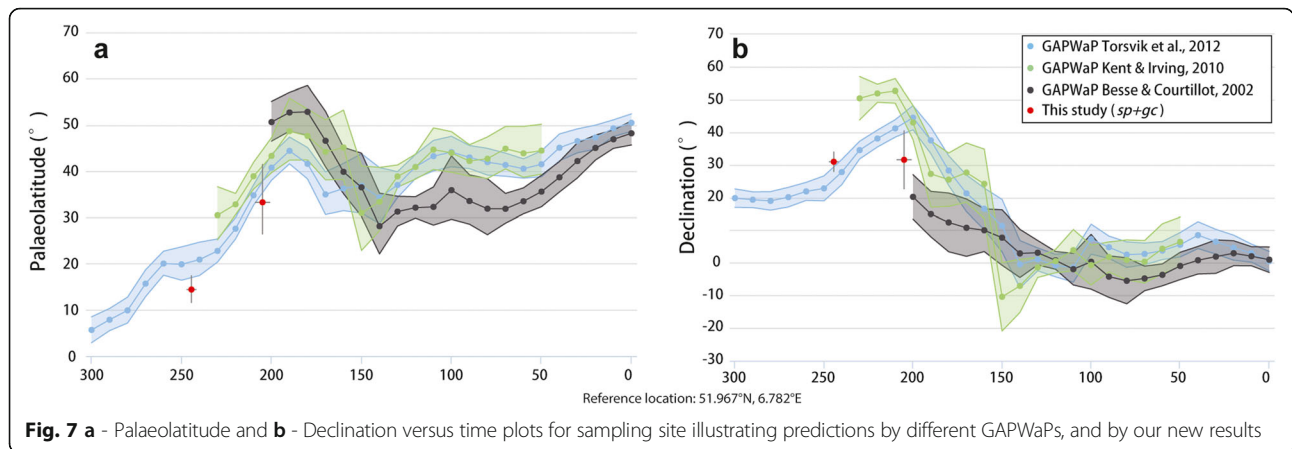
The palaeomagnetic data for the Anisian and the Rhaetian as observed in the Triassic sediments from the Winterswijk quarry enable us to calculate the corresponding palaeomagnetic poles (Table 1), which we define as the mean of the VGP distributions with its



corresponding  $K$  and  $A_{95}$ . The question is then which means —  $sp$  or  $sp+gc$  — we prefer. If we compare the two solutions for both the Anisian and the Rhaetian, we first note that the respective inclinations in both solutions are identical. The inclinations determine palaeolatitude and these are slightly lower but within an error of the predicted palaeolatitude for Eurasia according to the GAPWaP of Torsvik et al. (2012), which is based on 31 poles for the Anisian and 41 poles for the Rhaetian (Fig. 7). The Rhaetian palaeolatitude also conforms to the palaeolatitude predicted by the GAPWaP of Kent and Irving (2010), which does not continue back farther than 230 Ma and does not reach the Anisian (Fig. 7), and has a 210 Ma pole based on 11 palaeomagnetic poles. This further suggests that the Winterswijk sediments have faithfully recorded the field at the time of deposition, despite the pervasive overprint. We note that whilst

for limestones such as those of the Muschelkalk, inclination shallows due to compaction is uncommon, fine clastics like those of the Rhaetian may have experienced some compaction-induced flattening. Given a large number of great circle solutions, we cannot apply common correction methods like the E/I method (e.g., Tauxe et al. 2008). Applying a commonly used flattening factor of 0.6 (e.g., Torsvik et al. 2012) would lead to a corrected inclination for the Rhaetian of  $\sim 45^\circ$ , slightly higher than but still within an error of the GAPWaP of Torsvik et al. (2012).

In terms of declinations, however, there are some differences between the different solutions. For the Anisian, the declination from the  $sp+gc$  solution is close to the prediction for Eurasia by Kent and Irving (2010) and Torsvik et al. (2012) (Fig. 7) whereas from the  $sp$  solution the declination deviate by some  $15^\circ$  plots for the predicted declination curve, i.e., the observed  $sp$  declination is more



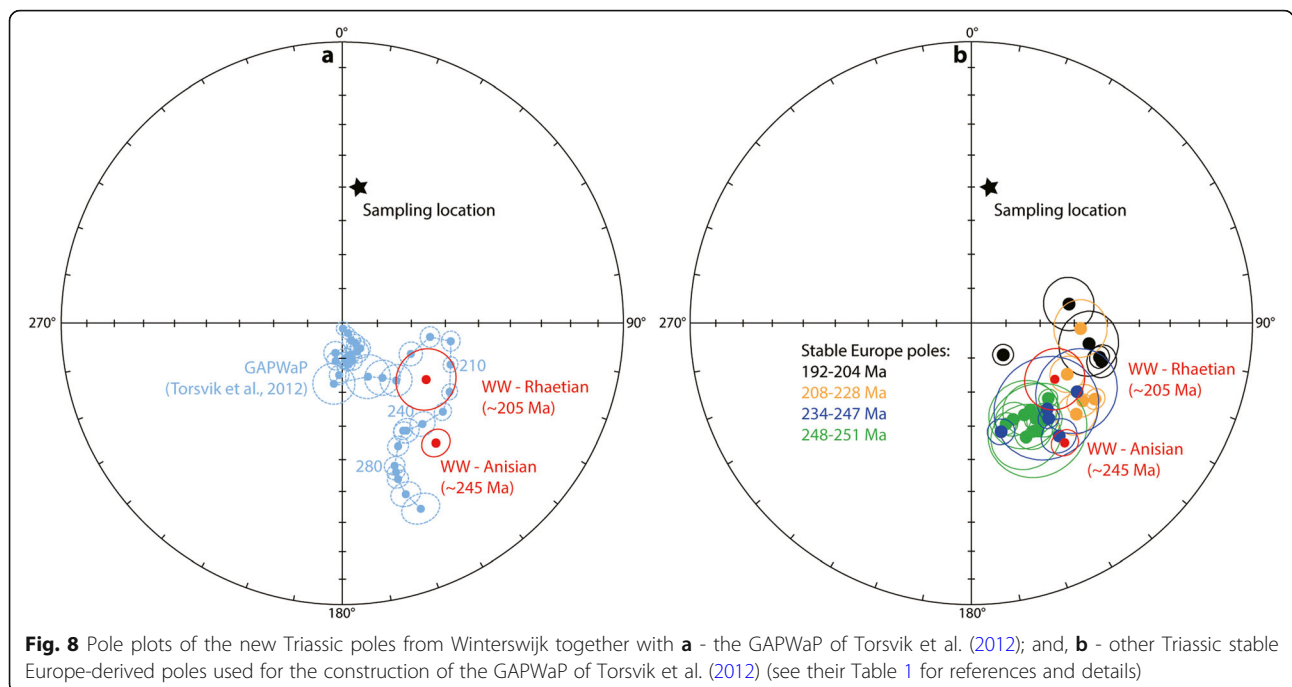
clockwise ( $\sim 41 \pm 8^\circ$ ) than expected ( $\sim 26 \pm 4^\circ$ ). For the Rhaetian, the *sp* solution has a declination ( $\sim 34 \pm 13^\circ$ ) that is close to the predicted declination ( $\sim 43^\circ$ ), whereas the *sp + gc* solution ( $\sim 32 \pm 9^\circ$ ) is lower than expected by Kent and Irving (2010) and Torsvik et al. (2012), although in line with the trend predicted by the GAPWaP of Besse and Courtillot (2002) that stops at 200 Ma.

The question then arises which of our solutions provides the best option. We note that solutions with so many great circles are not ideal, and as a rule, we prefer to have more setpoints than great circle solutions so that possible biases in great circle solutions do not dominate. In the Winterswijk case, however, the great circles are very distinct and consistent, and we used only great circle interpretations of well-defined overprint components and a clear tendency

towards the ChRM direction; components with noisy data in the Zijderveld diagrams were avoided. More importantly, the Rhaetian *sp* and *sp + gc* solutions share a common distribution and can be considered identical (Fig. 6). Hence, based on the consistent great circle solutions, the better fit of the Anisian *sp + gc* solution and the fact that the Rhaetian *sp* and *sp + gc* solutions are identical, we choose the *sp + gc* solutions as our best option to represent the Triassic poles of the Winterswijk sediments.

### 5.2 Triassic palaeogeography: position of the Netherlands in the heart of Pangea

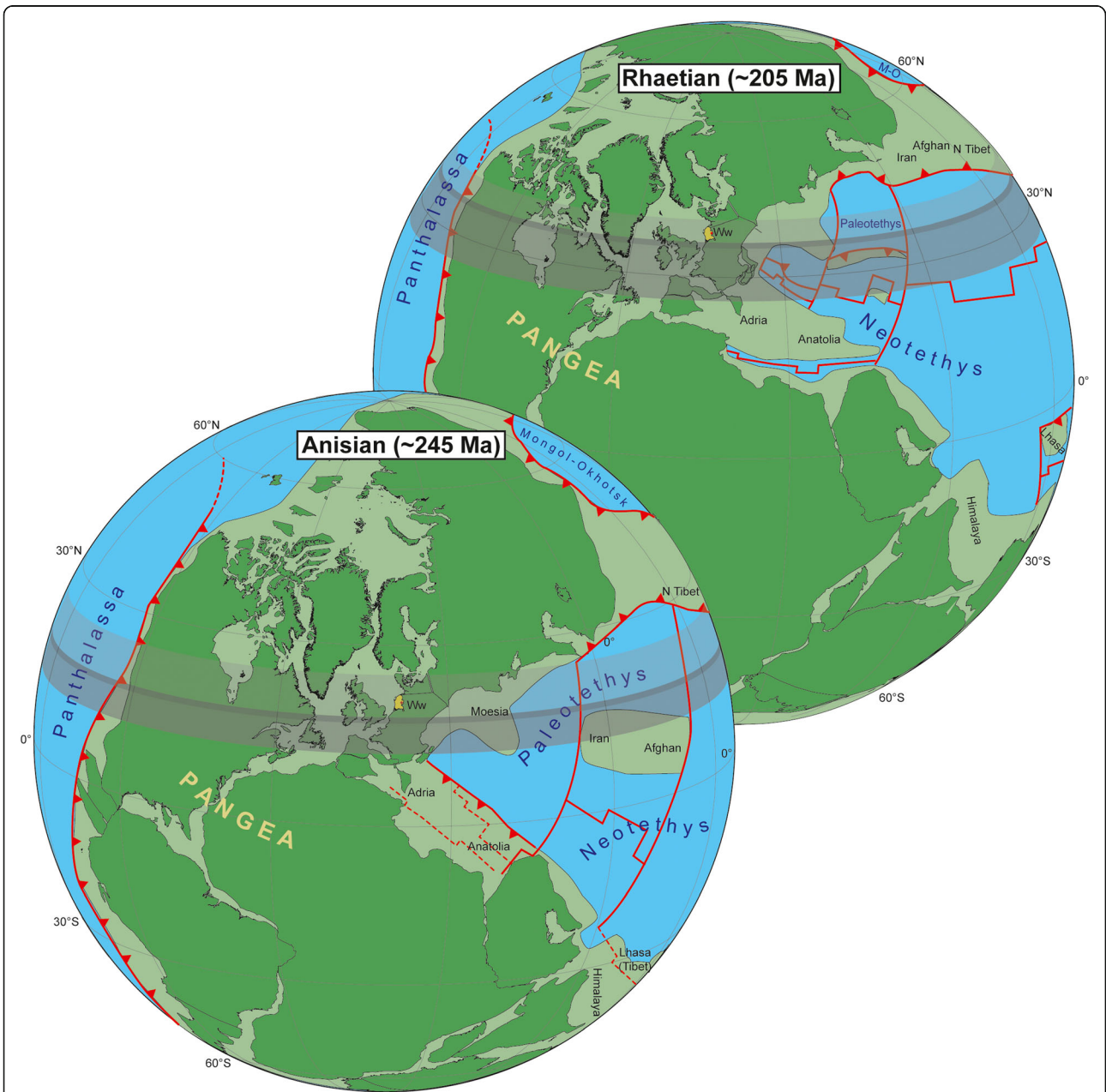
Our new palaeomagnetic pole from the Anisian of Winterswijk corresponds well with the GAPWaP of Torsvik et al. (2012) (Fig. 8a). It is important to realize that a



**Fig. 8** Pole plots of the new Triassic poles from Winterswijk together with **a** - the GAPWaP of Torsvik et al. (2012); and, **b** - other Triassic stable Europe-derived poles used for the construction of the GAPWaP of Torsvik et al. (2012) (see their Table 1 for references and details)

GAPWaP is based on several tens of poles per time interval, and that the error bars shown for the GAPWaP result from averaging these poles. The actual scatter of the poles is considerably larger than the average 95% confidence limit of the GAPWaP, and our data points fall within that scatter as illustrated in Fig. 8b, where our new data fall within the cluster of contemporaneous poles from stable

Eurasia. Figure 9 illustrates the palaeogeographic implication of these poles, showing the outline of the Netherlands within the supercontinent Pangea. This reconstruction follows Euler rotations for the major continents following Seton et al. (2012), with Neotethys and Paleotethys reconstructions following Muttoni et al. (2009), Gaina et al. (2013), Li et al. (2016) and Maffione and van Hinsbergen



**Fig. 9** Palaeogeographic maps of Pangea in the Anisian and the Rhaetian times, with the outline of the modern border of the Netherlands, and the sampling location in Winterswijk. Reconstructions based on Euler rotations of Seton et al. (2012) for the major continents, the Neotethys, and the Paleotethys based on Muttoni et al. (2009), Gaina et al. (2013), Li et al. (2016), Maffione and van Hinsbergen (2018), and East Asia following Van der Voo et al. (2015), cast in the palaeomagnetic reference frame of Torsvik et al. (2012). Light grey latitudinal band represents the palaeolatitudes predicted by the GAPWaP, darker grey band is the palaeolatitudes predicted by our new Winterswijk poles, and the darkest grey thin band represents the overlap between the two palaeolatitudes

(2018), and East Asia reconstruction following Van der Voo et al. (2015). The reconstruction is placed in the palaeomagnetic reference frame based on the GAPWaP of Torsvik et al. (2012), and illustrates the palaeolatitudes predicted by the Winterswijk sites.

Our new results further confirm that the Germanic Basin, with the large, epicontinental, shallow-marine Muschelkalk Sea, was located at low latitudes, at the Winterswijk locality  $\sim 15^\circ\text{N}$ , within the heart of Pangea. This latitude corresponds to the modern Sahara Desert, providing an analogue for the environment in which the Buntsandstein Formation was deposited; and corresponds to the modern Persian Gulf that may be an analogy to the palaeogeographic setting of the Muschelkalk Basin, although the Muschelkalk Basin was in a post-collisional environment while the Persian Gulf was in a foreland of an ongoing collision. During the Rhaetian, in the early stages of Pangea breakup, the Germanic Basin had shifted to more northerly latitudes of  $\sim 30^\circ\text{N}$ , corresponding to the modern latitude of e.g., Morocco and Sicily, causing the cooling trend on its way towards the modern latitude of  $52^\circ\text{N}$  (Table 1; Fig. 9).

## 6 Conclusions

We present two new Triassic palaeomagnetic poles for the stable Eurasian interior, based on the Anisian (247–242 Ma) blue-grey limestones and red marly limestones of the shallow-marine Muschelkalk Formation, and the Rhaetian (208–201 Ma) shallow-marine claystones exposed in a quarry at Winterswijk, the Netherlands. Rock magnetic experiments identify magnetite and hematite as carriers of the magnetization, depending on lithology. Demagnetization diagrams reveal a strong recent (GAD) overprint, as well as a second, low-inclination component at high coercivity and high unblocking temperature. This component is at times decaying towards the origin and is interpreted as the ChRM, whereas most specimens reveal great circles. Combining the great circle solutions with the isolated ChRMs, can lead to the two new palaeomagnetic poles that lie close to the predicted direction derived from GAPWaPs in Eurasian coordinates. Both poles may be included in future renditions of such paths, and confirm that the Germanic Basin, in which the Dutch Triassic was deposited, was located at an Anisian latitude around  $14 \pm 3^\circ\text{N}$ , and was characterized by a palaeogeography similar to that of the Persian Gulf today. By the Rhaetian time, the basin was at  $32 \pm 7^\circ\text{N}$  on its way to the modern latitude of  $52^\circ\text{N}$ .

## 7 Supplementary information

**Supplementary information** accompanies this paper at <https://doi.org/10.1186/s42501-019-0046-2>.

**Additional file 1: Supplementary Information 1a** Excel files of the IRM coercivity component fit for all samples. **Supplementary Information**

**1b** An overview of the input data for the IRM-CGL1.0 excel workbook (Kruiver et al. 2001).

**Additional file 2:** Demagnetization data (.dir files) and interpreted directions for statistical analysis (.pmag file) for the Winterswijk locality, compatible with the [Paleomagnetism.org](http://Paleomagnetism.org) online analysis software. WW 1–9 samples refer to the Anisian (Anisian\_pole.dir) whereby WW 1, 5, 6, 7 are from red marly limestones, and WW 2, 3, 4, 8, 9 are from grey limestones; WW 10 samples refer to the Rhaetian (Rhaetian\_pole.dir).

### Abbreviations

$^\circ\text{N}$ : Degrees North (latitude);  $A_{95}$ : 95% cone of confidence around the virtual geomagnetic poles; AF: Alternating field; ARM: Anhyseretic remanent magnetization; ChRM: Characteristic remanent magnetization; D: Declination; DP: Dispersion parameter; GAD: Geocentric axial dipole; GAP: Gradient acquisition plots; GAPWaP: Global apparent polar wander path;  $gc$ : Great circle; I: Inclination; IRM: Isothermal remanent magnetization; K (k or  $k$ ): Fisher (1953) precision parameter calculated on virtual geomagnetic poles; LAP: Linear acquisition plots; Lat: Latitude; Lon: Longitude; MAD: Maximum angular deviation; NRM: Natural remanent magnetization; nT, mT, T: nano-Tesla, milli-Tesla, Tesla; SAP: Standardized acquisition plots; SIRM: Saturation isothermal remanent magnetization;  $sp$ : Set point; VGP: Virtual geomagnetic pole;  $\Delta D_x$ : Declination error;  $\Delta I_x$ : Inclination error

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### Authors' contributions

LPPvH conceived of the study, performed sampling, and carried out palaeomagnetic laboratory analyses. DJJvH conceived of the study, performed sampling, analyzed data, and drafted the manuscript. CGL and MJD analyzed data and drafted the manuscript. BZ carried out palaeomagnetic laboratory analyses. MHL performed sampling and carried out palaeomagnetic laboratory analyses. All authors read and approved the final manuscript.

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### Availability of data and materials

All demagnetization diagrams, interpreted palaeomagnetic components, and statistical files are available in the Additional files 1 and 2 and can be uploaded, viewed, and edited on [Paleomagnetism.org](http://Paleomagnetism.org) (Koymans et al. 2016). Rock magnetic data files are available in the Additional files 1 and 2.

### Competing interests

The authors declare that they have no competing interests.

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